



Glasma

*– From the Color Glass Condensate
to a Quark-Gluon Plasma –*




– Gluon Production in pA and AA –

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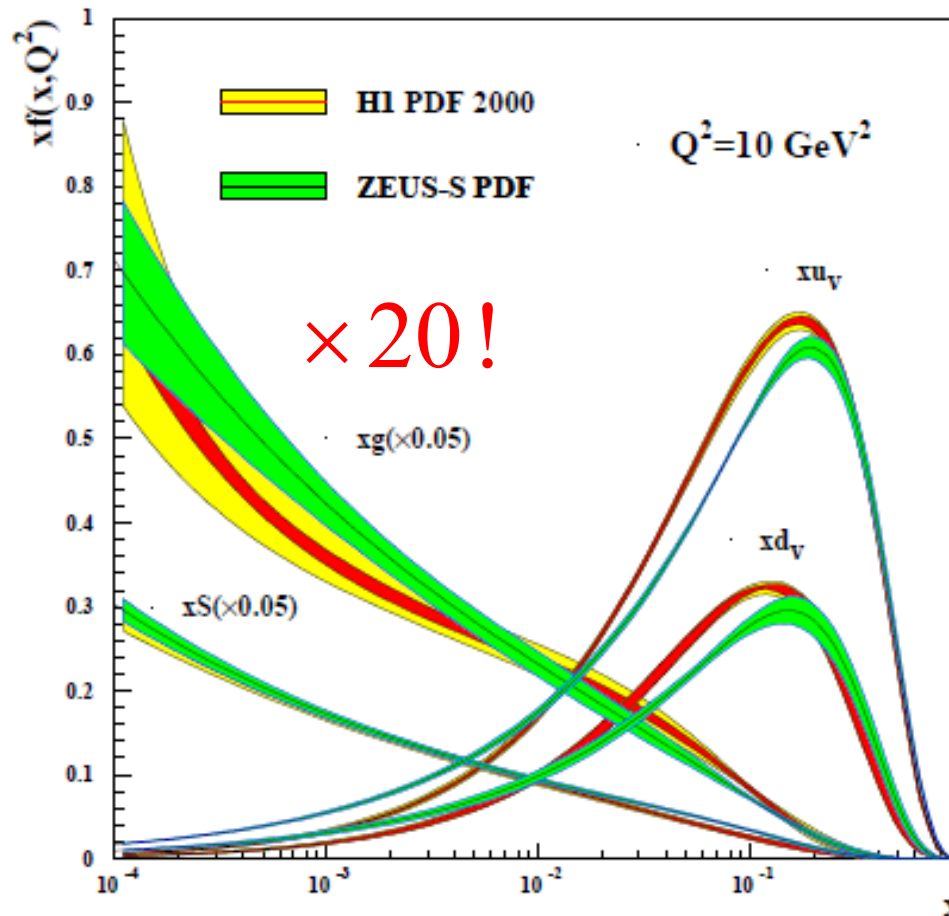
January 2009 at KEK

Talk Agenda

- 
- Saturation Model in the Small- x Region
 - Saturation scale as a function of x
 - McLerran-Venugopalan (MV) model
 - Particle Production in a Dilute-Dense Collision
 - One-gluon production
 - Two-gluon production
 - Particle Production in a Dense-Dense Collision
 - Gluon distribution \rightarrow Gluon production

HERA (ep collider) – x Evolution

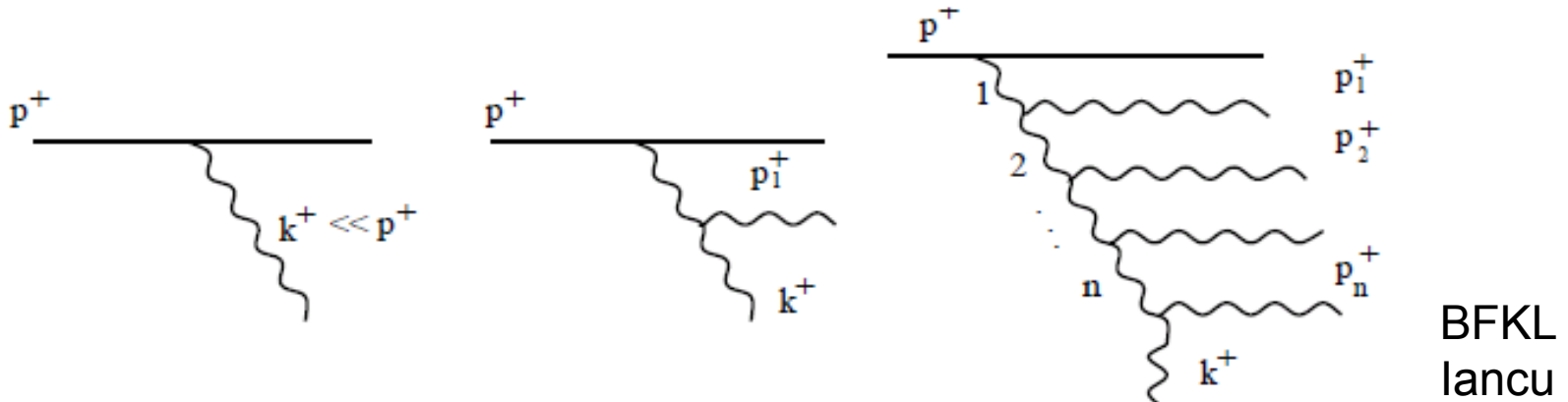
Quantum Evolution of PDFs



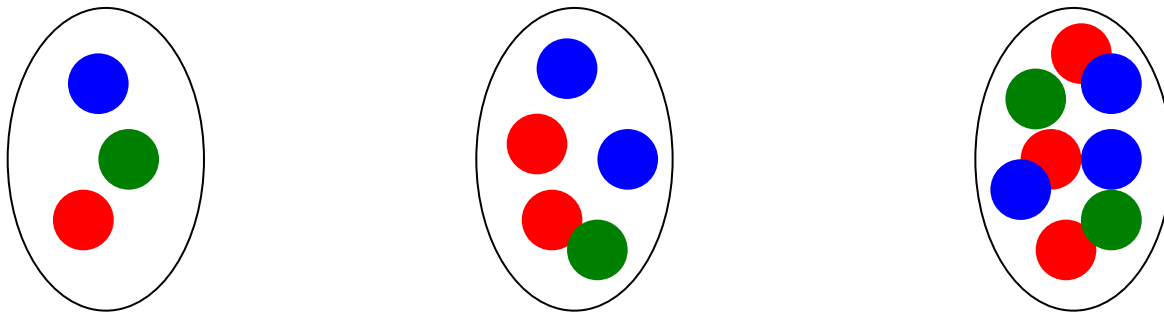
As x goes smaller than $\sim 10^{-2}$ gluon is dominant.

Going to Smaller x with Fixed Q^2

■ Gluon increases with a fixed transverse area



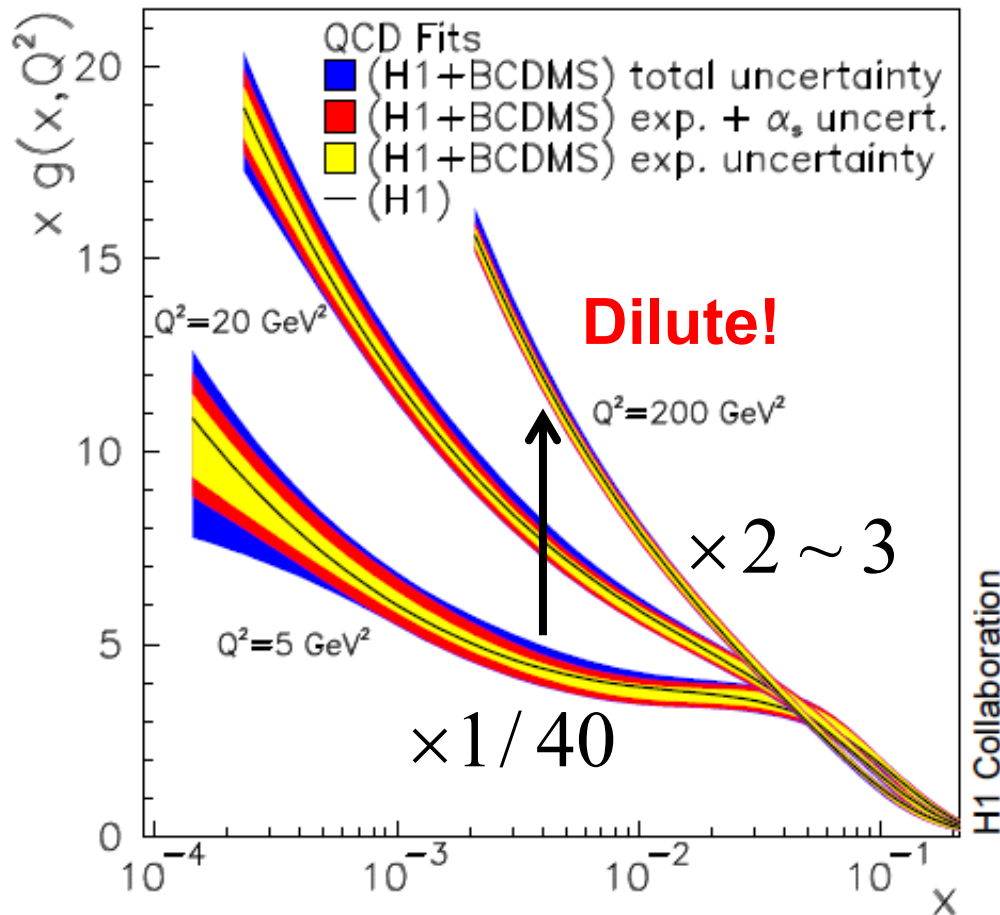
■ Graphically



small- x → Dense Gluon Matter

Q^2 Dependence

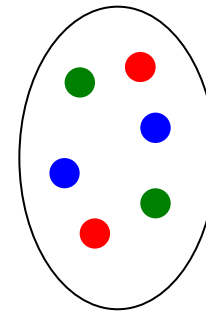
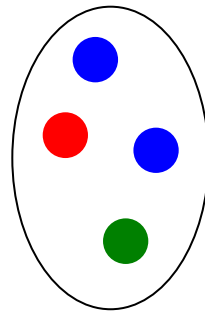
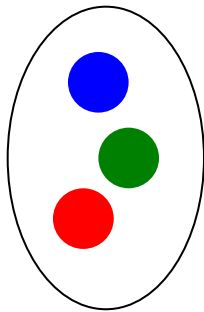
Quantum Evolution of PDFs



As Q^2 goes larger
gluon grows slowly.

Going to Larger Q^2 with Fixed x

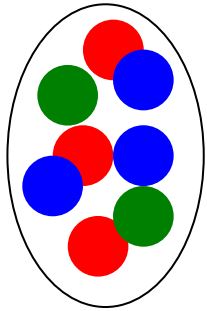
- Gluon slowly increases with a decreasing area
- Graphically, in the same way,



large $Q \rightarrow$ Dilute Gluon Matter

Saturation

- Gluons eventually cover the transverse area
at small x and small Q^2



$$\text{Area} \sim \pi R^2 \sim 75 \text{ GeV}^{-2} \text{ (proton)}$$

$$\text{Crammed density} \sim (N_c^2 - 1) \cdot Q^2 \cdot \pi R^2$$

- Condition for saturation: ~ 600 (for $Q^2 = 1 \text{ GeV}^2$)

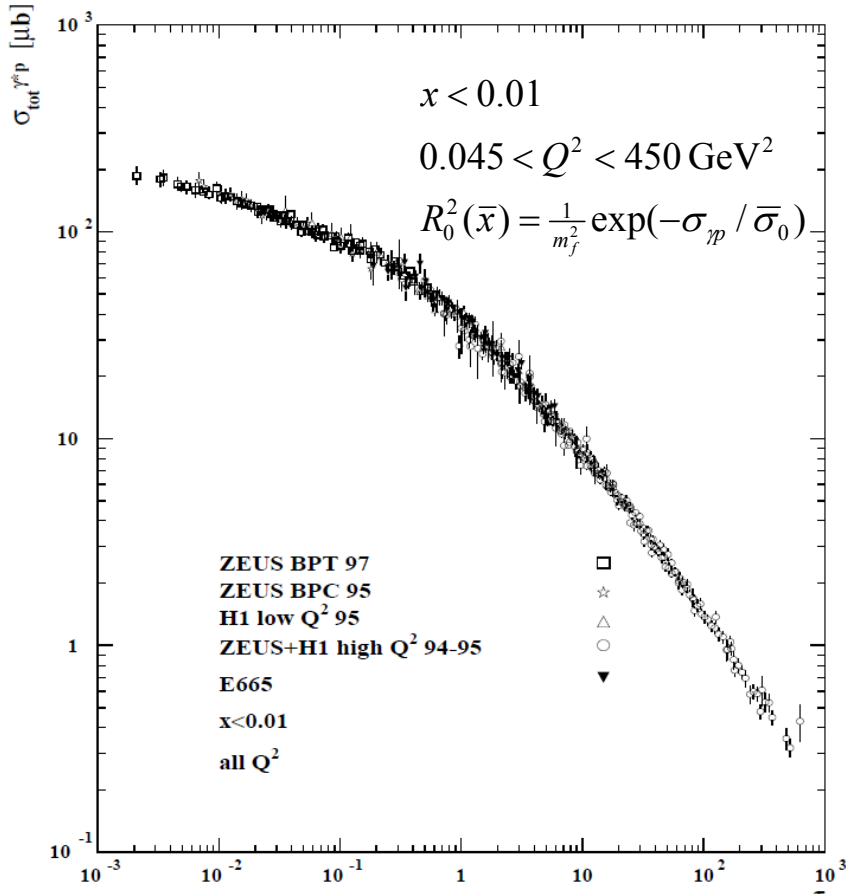
$$\frac{xg(x, Q_s)}{(N_c^2 - 1) \cdot Q_s^2 \cdot \pi R^2} \sim \frac{1}{\alpha_s N_c} \sim 1$$

Overlapping Factor

How can we expect the saturation effect in reality?

Scaling Behavior

Dipole Cross Section in a Saturation Model



$$\hat{\sigma}(x, r) = \sigma_0 \left(1 - \exp\left[-r^2 / 4R_0^2(x)\right] \right)$$

$$\sigma_{\gamma^*p}(x, Q^2) \rightarrow \sigma_{\gamma^*p}(Q^2 R_0^2(x))$$

Called the **"Geometric Scaling"**

$$R_0(x) = \frac{1}{Q_0} \left(\frac{x}{x_0} \right)^{\lambda/2}$$

Dias de Deus

$$Q_0 = 1 \text{ GeV}$$

$$x_0 = 3.04 \times 10^{-4}$$

$$\lambda = 0.288$$

lancu-Itakura-McLerran, lancu-Itakura-Munier

Stasto-Golec-Biernat-Kwiecinski Plot 2009 at KEK

Scaling Behavior Extended

■ Let us put some numbers... $x = 10^{-4} \rightarrow Q_s^2 = 1.38 \text{ GeV}^2$

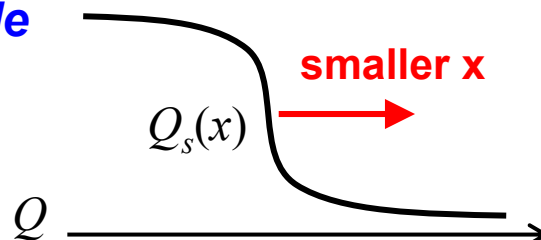
$$\frac{xg(x, Q)}{(N_c^2 - 1) \cdot Q^2 \cdot \pi R^2} \sim \frac{10}{8 \cdot 1.38 \cdot 75} \sim 0.01? \\ \times A^{1/3} (\sim 5.8 \text{ for Au})$$

■ Scaling is consistent with pQCD Iancu-Itakura-McLerran

"Extended Geometric Scaling"

■ BFKL (dilute regime) can fix the parameters

Dipole Amplitude



No need to realize saturation!

Message



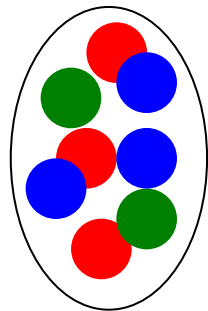
There are some examples in which
a saturation-model description can work well
even when saturation is not yet reached.

Saturation Model

■ Parametrized by a scaling variable

- $Q_s(x)$ encompasses the scaling property.
- Wave-function is characterized by $Q_s(x)$

■ McLerran-Venugopalan model

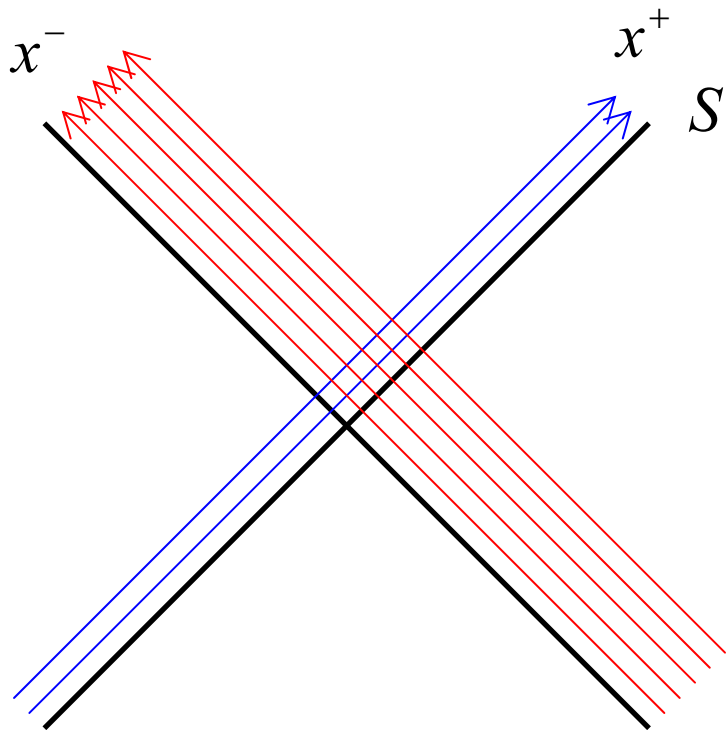


$$\mathcal{W}_x[\rho] = \exp \left[- \int d^3x \frac{|\rho(x)|^2}{2g^2 \mu_x^2} \right]$$

$$g\mu_x \sim Q_s(x)$$

Scattering Problem

Scattering Amplitude in the Eikonal Approx.



Dilute (p)

Dense (A)

$$S \sim \left\langle \left(\sum_{\{\rho_p\}} \mathcal{W}_x[\rho_p] \prod_{\{\rho_p\}} V \right) \left(\sum_{\{\rho_A\}} \mathcal{W}_{x'}[\rho_A] \prod_{\{\rho_A\}} W \right) \right\rangle$$

$$V(x_{\perp}) = \exp \left[ig \int dz^+ A^-(z^+, x_{\perp}) \right]$$

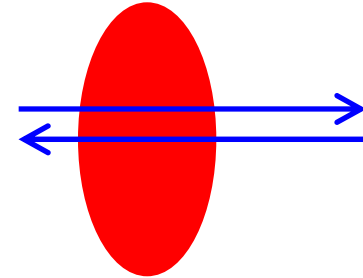
$$W(x_{\perp}) = \exp \left[ig \int dz^- A^+(z^-, x_{\perp}) \right]$$

In case of Dipole-CGC scattering

$$S \sim \left\langle \left\langle V(x_{\perp}) V^+(y_{\perp}) \right\rangle \right\rangle_{\rho_A}$$

Stationary-Point Approximation

Dipole Scattering Amplitude



$$\begin{aligned}
 S &\sim \left\langle \sum_{\{\rho_A\}} \mathcal{W}_x[\rho_A] \prod_{\{\rho_A\}} W \cdot V(x_\perp) V^+(y_\perp) \right\rangle \\
 &= \sum_{\{\rho_A\}} \mathcal{W}_x[\rho_A] \int_{p^- < xP^-} [\mathcal{D}A] V(x_\perp) V^+(y_\perp) \exp[iS_{\text{YM}}[A] + iS_{\text{source}}[\rho_A, W]] \\
 &= \left\langle \left\langle V(x_\perp) V^+(y_\perp) \right\rangle \right\rangle_{\rho_A}
 \end{aligned}$$

$$S_{\text{source}} = \frac{i}{gN_c} \int d^4x \text{tr}[\rho_A \ln W] \sim - \int d^4x \rho_A^a A_a^+$$

Easily solvable



$$\left. \frac{\partial S_{\text{YM}}}{\partial A_a^\mu} \right|_{A=\mathcal{A}} = \delta^{\mu-} \rho_A$$

Large enough $\rho_A \rightarrow$ Stationary-point approx.

Physical Observable

■ Dipole scattering amplitude

$$\begin{aligned} & \langle\langle V(x_{\perp})V^{+}(y_{\perp}) \rangle\rangle_{\rho_A} \\ &= \sum_{\{\rho_A\}} \mathcal{W}_x[\rho_A] \int_{p^+ < xP^+} [\mathcal{D}A] V(x_{\perp})V^{+}(y_{\perp}) \exp[iS_{\text{YM}}[A] + iS_{\text{source}}[\rho_A, W]] \\ &\sim \sum_{\{\rho_A\}} \mathcal{W}_x[\rho_A] V(x_{\perp})V^{+}(y_{\perp}) \Big|_{A=\mathcal{A}[\rho_A]} \quad (\text{function of } g\mu \sim Q_s) \end{aligned}$$

■ In general

$$\langle\langle O[A] \rangle\rangle_{\rho_A} \sim \int \mathcal{D}\rho_A \mathcal{W}_x[\rho_A] O[\mathcal{A}[\rho_A]]$$

**Quantum corrections
lead to $\mathcal{W}_x \rightarrow \mathcal{W}_{x+\delta x}$
i.e. small-x evolution**

Classical Solution



$$x^- \quad \alpha_i^{(2)}(x_\perp)$$

$$\mathcal{A}^+ = \mathcal{A}^- = 0$$

$$\mathcal{A}_i = \alpha_i^{(2)} - \frac{1}{ig} W(x_\perp) \partial_i W^+(x_\perp)$$

$$W^+(x_\perp) = P \exp \left[-ig \int dz^+ \frac{1}{\partial_\perp^2} \rho_A(x_\perp) \delta(z^+) \right]$$

static
(time dilation)

$$\delta^{\mu-} \delta(x^+) \rho_A(x_\perp)$$

thin
(Lorentz contract)

c.f. in EM

$$\partial_\perp^2 \phi = -\rho \quad (\text{Gauss law})$$

$$\rightarrow \phi' = 0$$

$$\rightarrow A_i = \frac{1}{ie} e^{ie\phi} \partial_i e^{-ie\phi} \quad (= -\partial_i \phi)$$

One-Gluon Production in pA

LSZ reduction formula

$$\langle \mathbf{p}; \lambda, a | \Omega \rangle = i \epsilon_{\mu}^{(\lambda)}(\mathbf{p}) \int d^4x e^{ip \cdot x} \square_x \langle \Omega | A_a^{\mu}(x) | \Omega \rangle$$

In LC gauge

$$\frac{dN_g}{dy} = \frac{1}{4\pi} \int \frac{d^2 \mathbf{p}_{\perp}}{(2\pi)^2} p^2 \mathcal{A}_a^{i*}(p) p^2 \mathcal{A}_a^i(p)$$



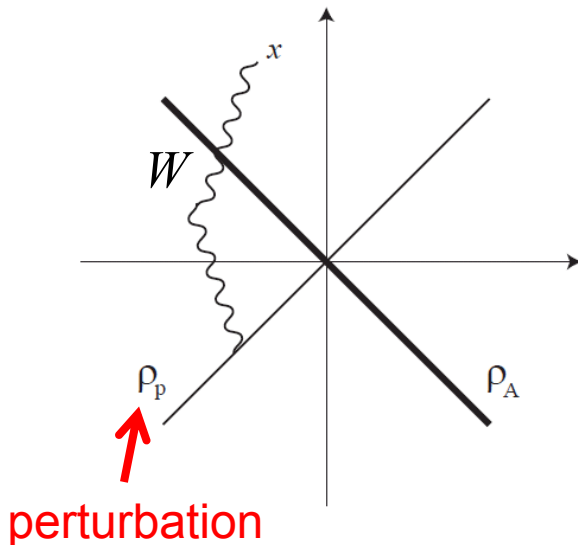
Final expression is given in terms of

$$\langle \rho_p \rho_p \rangle \rightarrow \varphi_p \quad \langle WW \rangle \rightarrow \varphi_A \quad K_t\text{-factorization}$$

which is calculable as a function of $g\mu$
(higher-twist)

Dumitru-McLerran, Blaizot-Gelis-Venugopalan

January 2009 at KEK



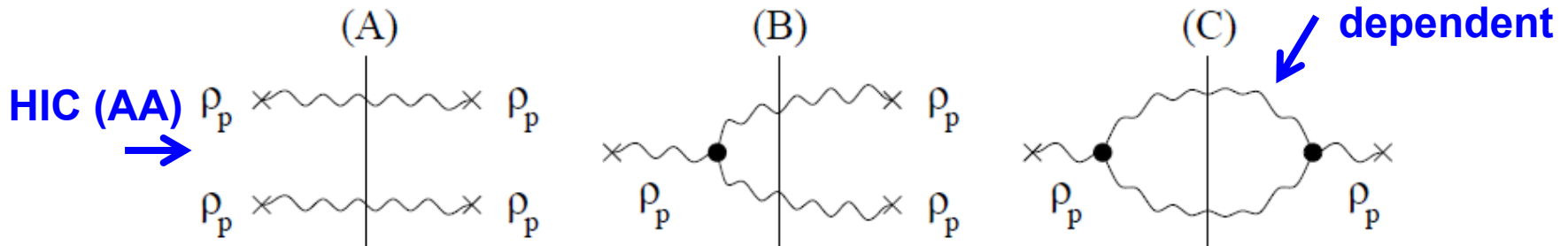
Two-Gluon Production in pA

LSZ reduction formula

$$\begin{aligned} & \langle \mathbf{p}, \lambda, a; \mathbf{q}, \sigma, b | \Omega \rangle \\ &= i \epsilon_{\mu}^{(\lambda)}(\mathbf{p}) \int d^4x e^{ip \cdot x} \square_x i \epsilon_{\nu}^{(\sigma)}(\mathbf{q}) \int d^4y e^{iq \cdot y} \square_y \langle \Omega | T[A_a^{\mu}(x) A_b^{\nu}(y)] | \Omega \rangle \end{aligned}$$

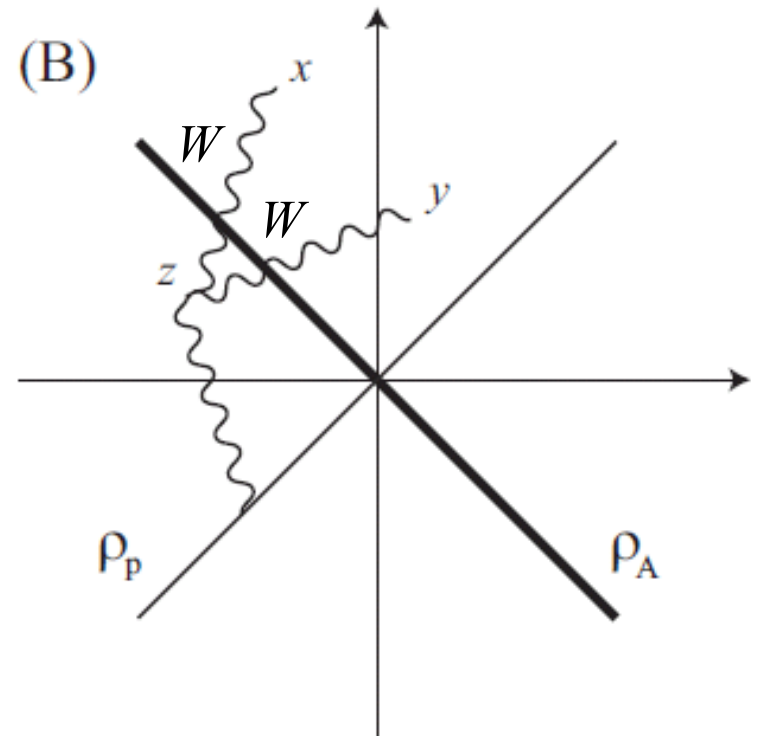
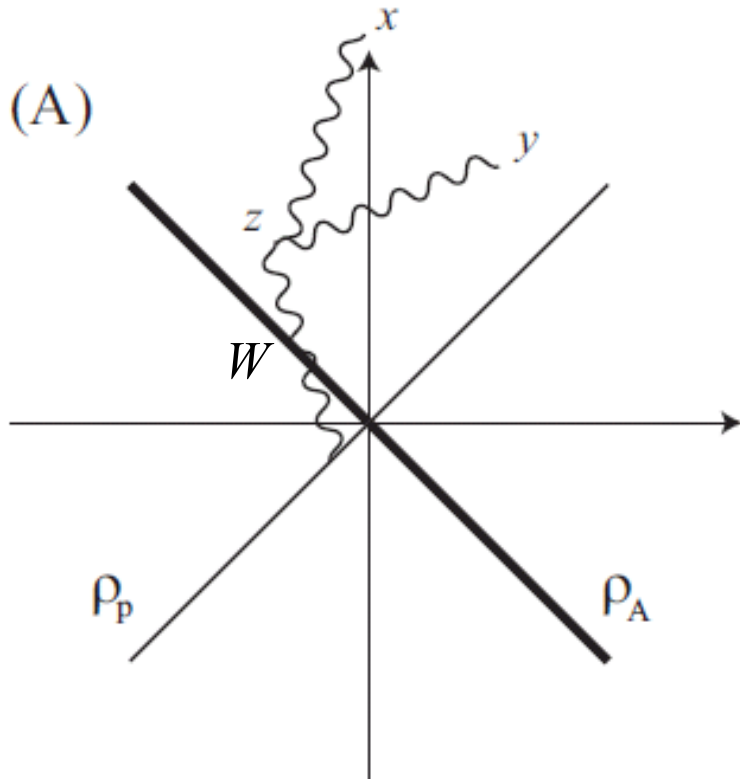
In LC gauge

$$\begin{aligned} \frac{dN_{gg}}{dy_1 dy_2} &= \frac{1}{16\pi^2} \int \frac{d^2\mathbf{p}_{\perp}}{(2\pi)^2} \frac{d^2\mathbf{q}_{\perp}}{(2\pi)^2} \sum_{\lambda, \sigma, a, b} \left\langle \left| \langle \mathbf{p}, \lambda, a; \mathbf{q}, \sigma, b | \Omega \rangle \right|^2 \right\rangle \\ &= \frac{dN_g}{dy_1} \frac{dN_g}{dy_2} + \frac{1}{16\pi^2} \int \frac{d^2\mathbf{p}_{\perp}}{(2\pi)^2} \frac{d^2\mathbf{q}_{\perp}}{(2\pi)^2} (p^2)^2 (q^2)^2 2\text{Re} \left[\mathcal{A}_a^{*i}(p) \mathcal{G}_{ab}^{ij}(p, q) \mathcal{A}_b^{*j}(q) \right] \\ &+ \frac{1}{16\pi^2} \int \frac{d^2\mathbf{p}_{\perp}}{(2\pi)^2} \frac{d^2\mathbf{q}_{\perp}}{(2\pi)^2} (p^2)^2 (q^2)^2 \mathcal{G}_{ba}^{\dagger ji}(q, p) \mathcal{G}_{ab}^{ij}(p, q). \end{aligned}$$



Two-Gluon Production in pA

Diagrams for the Connected Part



Fukushima-Hidaka

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Dense-Dense (AA) Collision

Glasma = Glass + Plasma

On the light cone ($\tau = 0$)

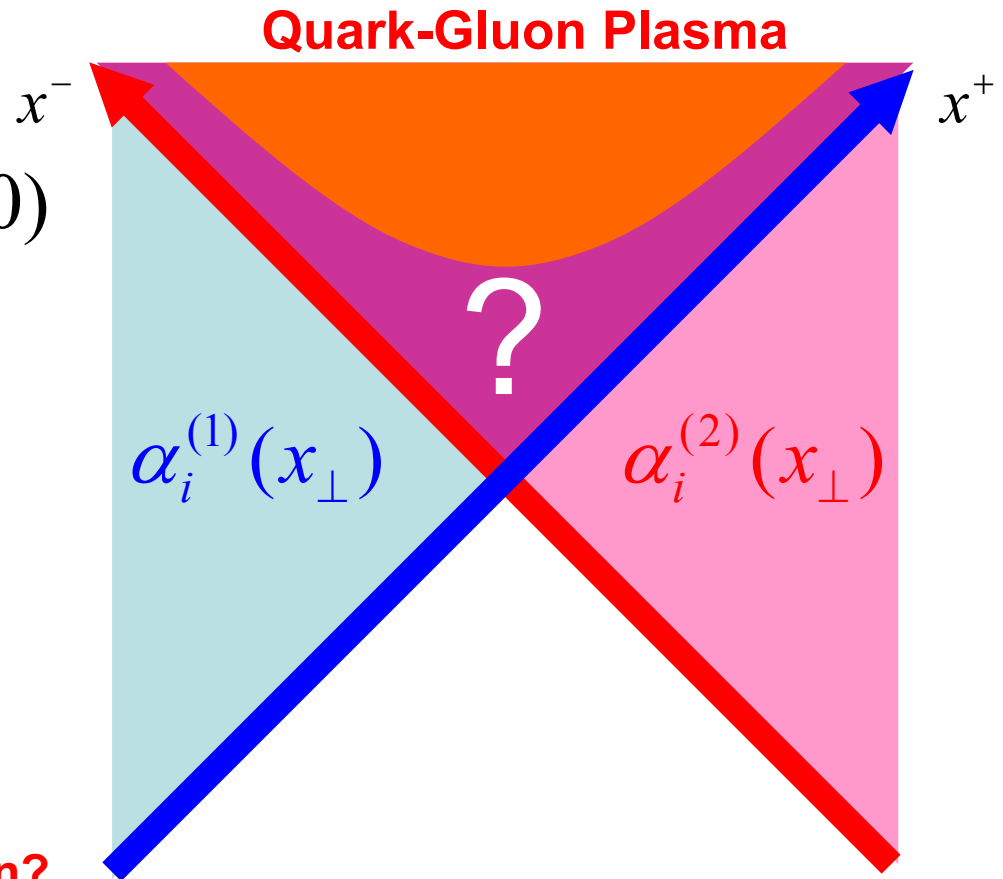
$$\mathcal{A}_i = \alpha_i^{(1)} + \alpha_i^{(2)}$$

$$\mathcal{A}_\eta = 0$$

$$\mathcal{E}^i = 0$$

$$\mathcal{E}^\eta = ig[\alpha_i^{(1)}, \alpha_i^{(2)}]$$

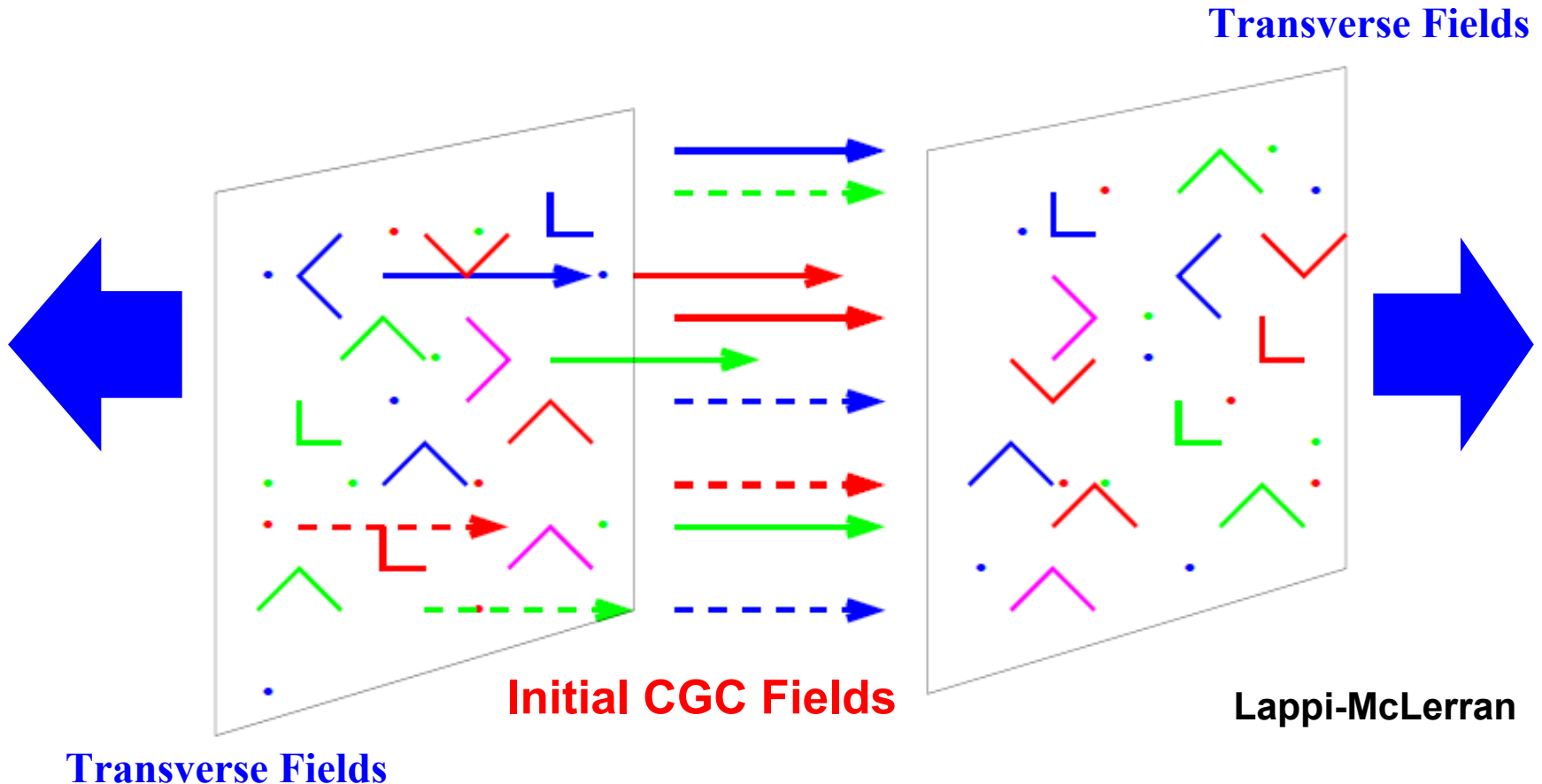
How to compute gluon production?



Intuitive Picture at $\tau = 0$



Longitudinal Fields between Nucleus Sheets



Gaussian Average at τ^0

■ Energy density

$$\varepsilon_{(0)} = \left\langle \text{tr} [E_{(0)}^\eta E_{(0)}^\eta + B_{(0)}^\eta B_{(0)}^\eta] \right\rangle$$

$$= 2N_c(N_c^2 - 1)g^2 \langle \alpha\alpha \rangle^2 = g^6 \mu_A^4 \cdot \frac{3}{\pi^2} \left[\ln \frac{\Lambda}{m} \right]^2$$

UV div.
(unphysical)

IR div.
(physical)

■ Momentum decomposition

$$a \equiv \frac{N_c (g^2 \mu_A)^2}{2m^2}$$

$$\varepsilon_{(0)}(k_\perp) = \frac{1}{V} \int d^2 \mathbf{x}_\perp d^2 \mathbf{y}_\perp e^{-i\mathbf{k}_\perp(\mathbf{x}_\perp - \mathbf{y}_\perp)}$$

$$\times \left\langle \text{tr} [E_{(0)}^\eta(\mathbf{x}_\perp) E_{(0)}^\eta(\mathbf{y}_\perp) + B_{(0)}^\eta(\mathbf{x}_\perp) B_{(0)}^\eta(\mathbf{y}_\perp)] \right\rangle$$

$$= \frac{1}{2} N_c (N_c^2 - 1) g^6 \mu_A^4$$

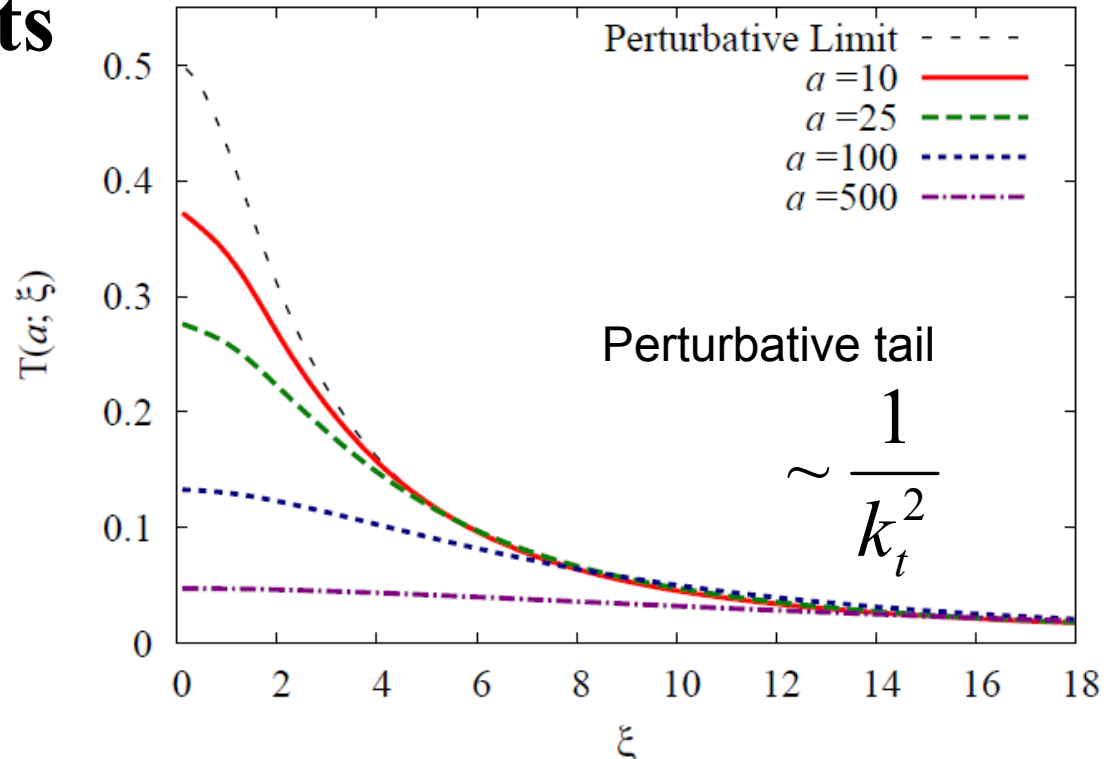
Expressed by $K_1(x)$

$$\times \int \frac{d^2 \mathbf{q}_{1\perp}}{(2\pi)^2} \frac{d^2 \mathbf{q}_{2\perp}}{(2\pi)^2} \frac{d^2 \mathbf{q}_{3\perp}}{(2\pi)^2} \frac{\bar{C}_{\text{adj}}(\mathbf{q}_{1\perp}) \bar{C}_{\text{adj}}(\mathbf{q}_{2\perp})}{(\mathbf{q}_{3\perp}^2 + m^2) [(k'_\perp - \mathbf{q}_{3\perp})^2 + m^2]}$$

Spectral Intensity

Numerical Results

Physical choice is
 $a = 10 \sim 40$ at RHIC



$$\varepsilon_{(0)}(k_{\perp}) = \frac{1}{4\pi m^2} N_c (N_c^2 - 1) g^6 \mu_A^4 \mathcal{T}(a; k_{\perp}/m)$$

UV Dominance

UV Contribution Dominance

- UV divergences come from derivatives.
- Time evolution from derivative terms in EoM.

$$E^\eta(\tau, \mathbf{k}_\perp) = E_{(0)}^\eta(\mathbf{k}_\perp) J_0(k_\perp \tau),$$

$$E^i(\tau, \mathbf{k}_\perp) = E_{(2)}^i(\mathbf{k}_\perp) \frac{2\tau}{k_\perp} J_1(k_\perp \tau),$$

$$B^\eta(\tau, \mathbf{k}_\perp) = B_{(0)}^\eta(\mathbf{k}_\perp) J_0(k_\perp \tau),$$

$$B^i(\tau, \mathbf{k}_\perp) = B_{(2)}^i(\mathbf{k}_\perp) \frac{2\tau}{k_\perp} J_1(k_\perp \tau).$$

**Decreasing functions
as k goes larger.**

**UV div is tamed by
expansion.**

Gluon Distribution

Energy and Gluon Distribution (Multiplicity)

$$\varepsilon = \langle \mathcal{H} \rangle \quad n(\tau, \mathbf{k}_\perp) = \frac{1}{\sqrt{k_\perp^2 + m^2}} \varepsilon(\tau, \mathbf{k}_\perp)$$

Fourier decomposition

$$\frac{dE(\tau)}{d\eta} \quad \frac{dN(\tau)}{d\eta}$$

$$\begin{aligned} \frac{dN(\tau)}{d\eta} &= \pi R_A^2 \tau n(\tau) \\ &= \frac{3\pi R_A^2}{\pi^2 m^2} \cdot \frac{1}{g^2} (g^2 \mu_A)^4 (m\tau) I_N(a; m\tau) \end{aligned}$$

Result

Fujii-Fukushima-
-Hidaka

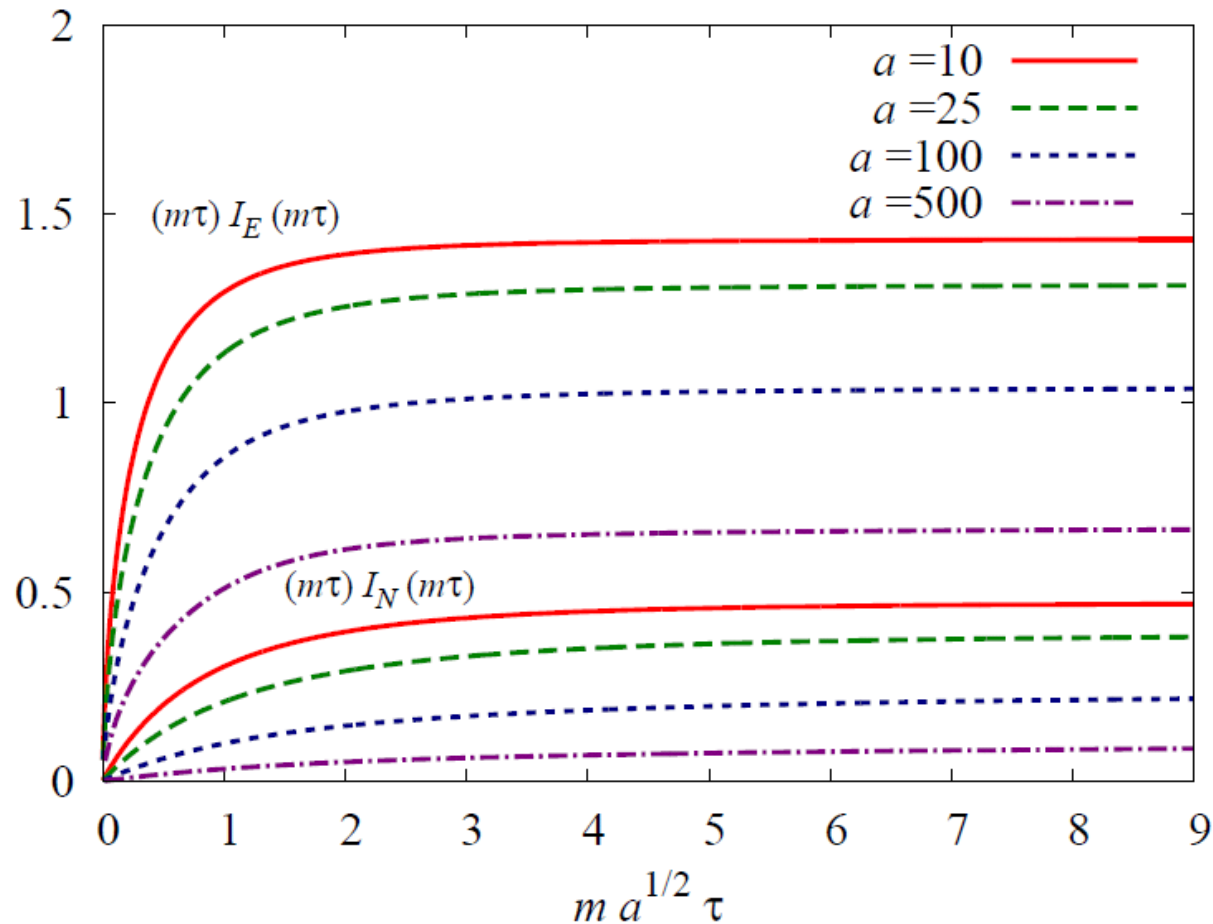
$$I_N(a; m\tau) = \int_0^\infty d\xi \frac{\xi \mathcal{T}(\xi)}{\sqrt{\xi^2 + 1}} \left\{ [J_0(\xi m\tau)]^2 + [J_1(\xi m\tau)]^2 \right\}$$

c.f. Gunion-Bertsch

Consistent with the LSZ reduction formula! (Gelis-Venugopalan)

Well-Defined Asymptotics

Numerical Results ($\sim dN/d\eta$)



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Numerics


Results (with non-linear corrections) for $Q_s^2 = 1 \sim 2 \text{ GeV}^2$

a	$dE/d\eta$ (GeV)	$dN/d\eta$	c
10	$(2.8 - 7.8) \times 10^3$	$(0.7 - 1.4) \times 10^3$	0.86
25	$(2.5 - 7.0) \times 10^3$	$(0.9 - 1.8) \times 10^3$	1.10
100	$(1.6 - 4.5) \times 10^3$	$(1.0 - 1.9) \times 10^3$	1.18
500	$(1.0 - 2.9) \times 10^3$	$(1.0 - 1.9) \times 10^3$	1.17

$$\frac{1}{\pi R_A^2} \frac{dN}{d\eta} = c \frac{N_c^2 - 1}{\pi g^2 N_c} Q_s^2$$

**Consistent with the empirical value at $a = 25$ and $Q_s^2 = 1 \text{ GeV}^2$
Hadron multiplicity at mid rapidity ~ 1100**

Summary

- 
- Saturation models enable us to carry out economical resummation of multiple scattering with dense color source.
 - Gluon multiplicity can be computed.
 - One-gluon production in pA
 - Two-gluon production in pA, ...etc
 - Gluon multiplicity in the expanding geometry in AA can be computed.