New Applications of Resummation in Non-abelian Gauge Theories: QED SQCD Exponentiation for LHC Physics and Resummed Quantum Gravity

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#### Outline:

- Introduction
- Review of YFS Theory and Its Extension to QCD
- Extension to QED (QCD and Quantum Gravity
- QED © QCD Threshold Corrections, Shower/ME Matching and IR-Improved DGLAP Theory at the LHC
- Final State of Hawking Radiation
- Conclusions

Papers byB.F.L.W., S. Jadach and B.F.L. Ward, S. Jadach, *et al.*,B.F.L.W. and S. Yost, M. Phys. Lett. A **14** (1999) 491, hep-ph/0205062; *ibid.* **12** (1997) 2425; *ibid.***19** (2004) 2113;

hep-ph/0503189,0508140,0509003

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 FNAL/RHIC tī PRODUCTION; POLARIZED pp PROCESSES; bb̄ PRODUCTION; J/Ψ PRODUCTION: SOFT n(G) EFFECTS ALREADY NEEDED

 $\Delta m_t = 5.1~{
m GeV}$  with SOFT n(G) UNCERTAINTY  $\sim$  2-3 GeV, ..., ETC.

- FOR THE LHC/TESLA/LC, THE REQUIREMENTS WILL BE EVEN MORE DEMANDING AND OUR QCD SOFT n(G) MC EXPONENTIATION RESULTS WILL BE AN IMPORTANT PART OF THE NECESSARY THEORY – YFS EXPONENTIATED  $\mathcal{O}(\alpha_s^2)L$ , IN THE PRESENCE OF SHOWERS, ON AN EVENT-BY-EVENT BASIS, WITHOUT DOUBLE COUNTING AND WITH EXACT PHASE SPACE.
- HOW RELEVANT ARE QED HIGHER ORDER CORRECTIONS WHEN QCD IS CONTROLLED AT  $\sim 1\%$  PRECISION?
- CROSS CHECK OF QCD LITERATURE:
  1. PHASE SPACE CATANI, CATANI-SEYMOUR, ALL INITIAL PARTONS MASSLESS
  2. RESUMMATION – STERMAN, CATANI ET AL., BERGER ET AL., ....
  3. NO-GO THEOREMS
  - 4. IR QCD EFFECTS IN DGLAP THEORY

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• CROSS CHECK OF QED LITERATURE:

 ESTIMATES BY SPIESBERGER, STIRLING, ROTH and WEINZIERL – FEW PER MILLE EFFECTS FROM QED CORRECTIONS TO STR. FN. EVOLUTION.
 WELL-KNOWN POSSIBLE ENHANCEMENT OF QED CORRECTIONS AT THRESHOLD, ESPECIALLY IN RESONANCE PRODUCTION

 $\Rightarrow$  HOW BIG ARE THESE EFFECTS AT THE LHC?

- TREAT QED AND QCD SIMULTANEOUSLY IN THE YFS EXPONENTIATION TO ESTIMATE THE ROLE OF THE QED AND TO ILLUSTRATE AN APPROACH TO SHOWER/ME MATCHING.
- QUANTUM GENERAL RELATIVITY:STILL NO
   PHENOMENOLOGICALLY TESTED THEORY
- OUTSTANDING ISSUES: FINAL STATE OF HAWKING RADIATION, ... – FERTILE GROUND FOR RESUMMATION;
   SEE ALSO WORK BY REUTER ET AL., LITIM, DONOGHUE ET AL., CAVAGLIA, SOLA ET AL., ETC.

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Review of YFS Theory and Its Extension to QCD

QED CASE – S. Jadach et al., YFS2, YFS3, BHLUMI, BHWIDE, KORALZ, KKMC, YFSWW3, YFSZZ, KoralW For  $e^+(p_1)e^-(q_1) \rightarrow \overline{f}(p_2)f(q_2) + n(\gamma)(k_1, \cdot, k_n)$ , renormalization group improved YFS theory (PRD36(1987)939) gives

$$d\sigma_{exp} = e^{2\alpha \operatorname{Re} B + 2\alpha \tilde{B}} \sum_{n=0}^{\infty} \frac{1}{n!} \int \prod_{j=1}^{n} \frac{d^{3}k_{j}}{k_{j}^{0}} \int \frac{d^{4}y}{(2\pi)^{4}} e^{iy(p_{1}+q_{1}-p_{2}-q_{2}-\sum_{j}k_{j})+D}$$
$$\bar{\beta}_{n}(k_{1},\dots,k_{n}) \frac{d^{3}p_{2}d^{3}q_{2}}{p_{2}^{0}q_{2}^{0}}$$

where the YFS real infrared function  $\tilde{B}$  and the virtual infrared function B are known and where we note the usual connections

$$2\alpha \,\tilde{B} = \int^{k \le K_{max}} \frac{d^3k}{k_0} \tilde{S}(k)$$

$$D = \int d^3k \frac{\tilde{S}(k)}{k^0} \left( e^{-iy \cdot k} - \theta (K_{max} - k) \right)$$
<sup>(2)</sup>

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for the standard YFS infrared emission factor

$$\tilde{S}(k) = \frac{\alpha}{4\pi^2} \left[ Q_f Q_{(\bar{f})'} \left( \frac{p_1}{p_1 \cdot k} - \frac{q_1}{q_1 \cdot k} \right)^2 + (\dots) \right]$$
(3)

if  $Q_f$  is the electric charge of f in units of the positron charge. For example, the YFS hard photon residuals  $\bar{\beta}_i$  in (1), i = 0, 1, 2, are given in S. Jadach *et al.*,CPC102(1997)229 for BHLUMI 4.04  $\Rightarrow$  YFS exponentiated exact  $\mathcal{O}(\alpha)$  and LL  $\mathcal{O}(\alpha^2)$  cross section for Bhabha scattering via a corresponding Monte Carlo realization of (1).



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In hep-ph/0210357(ICHEP02), Acta Phys.Polon.B33,1543-1558,2002, we have extended the YFS theory to QCD:

$$d\hat{\sigma}_{\exp} = \sum_{n} d\hat{\sigma}^{n}$$

$$= e^{\text{SUM}_{\text{IR}}(\text{QCD})} \sum_{n=0}^{\infty} \int \prod_{j=1}^{n} \frac{d^{3}k_{j}}{k_{j}} \int \frac{d^{4}y}{(2\pi)^{4}} e^{iy \cdot (P_{1}+P_{2}-Q_{1}-Q_{2}-\sum k_{j})+D_{\text{QCD}}}$$

$$* \tilde{\bar{\beta}}_{n}(k_{1}, \dots, k_{n}) \frac{d^{3}P_{2}}{P_{2}^{0}} \frac{d^{3}Q_{2}}{Q_{2}^{0}}$$
(4)

where now the hard gluon residuals  $ilde{areta}_n(k_1,\ldots,k_n)$  defined by

$$\tilde{\bar{\beta}}_n(k_1,\ldots,k_n) = \sum_{\ell=0}^{\infty} \tilde{\bar{\beta}}_n^{(\ell)}(k_1,\ldots,k_n)$$

are free of all infrared divergences to all orders in  $lpha_s(Q)$ .

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- We stress that the arguments in the earlier papers (DeLaney *et al.* PRD52(1995)108, PLB342(1995)239) are not really sufficient to derive the respective analog of eq.(4); for, they did not really expose the compensation between the left over genuine non-Abelian IR virtual and real singularities between  $\int dPh\bar{\beta}_n$  and  $\int dPh\bar{\beta}_{n+1}$  respectively that really allows us to isolate  $\tilde{\beta}_j$  and distinguishes QCD from QED, where no such compensation occurs.
- Our exponential factor corresponds to the N = 1 term in the exponent in Gatheral's formula (Phys. Lett.B133(1983)90) for the general exponentiation of the eikonal cross sections for non-Abelian gauge theory; his result is an approximate one in which everything that does not eikonalize and exponentiate is dropped whereas our result (4) is exact.

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Extension to QED QCD and Quantum Gravity

Simultaneous exponentiation of QED and QCD higher order effects, hep-ph/0404087,

gives

$$B_{QCD}^{nls} \to B_{QCD}^{nls} + B_{QED}^{nls} \equiv B_{QCED}^{nls},$$
  

$$\tilde{B}_{QCD}^{nls} \to \tilde{B}_{QCD}^{nls} + \tilde{B}_{QED}^{nls} \equiv \tilde{B}_{QCED}^{nls},$$
  

$$\tilde{S}_{QCD}^{nls} \to \tilde{S}_{QCD}^{nls} + \tilde{S}_{QED}^{nls} \equiv \tilde{S}_{QCED}^{nls}$$

which leads to

$$d\hat{\sigma}_{\exp} = e^{\text{SUM}_{\text{IR}}(\text{QCED})} \sum_{n,m=0}^{\infty} \int \prod_{j_1=1}^{n} \frac{d^3 k_{j_1}}{k_{j_1}}$$
$$\prod_{j_2=1}^{m} \frac{d^3 k'_{j_2}}{k'_{j_2}} \int \frac{d^4 y}{(2\pi)^4} e^{iy \cdot (p_1 + q_1 - p_2 - q_2 - \sum k_{j_1} - \sum k'_{j_2}) + D_{\text{QCED}}}$$
$$\tilde{\bar{\beta}}_{n,m}(k_1, \dots, k_n; k'_1, \dots, k'_m) \frac{d^3 p_2}{p_2^0} \frac{d^3 q_2}{q_2^0},$$

where the new YFS residuals

 $ilde{eta}_{n,m}(k_1,\ldots,k_n;k_1',\ldots,k_m')$ , with n hard gluons and m hard photons,

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(5)

(6)

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represent the successive application of the YFS expansion first for QCD and subsequently for QED.

The infrared functions are now

$$SUM_{IR}(QCED) = 2\alpha_s \Re B_{QCED}^{nls} + 2\alpha_s \tilde{B}_{QCED}^{nls}$$

$$D_{QCED} = \int \frac{dk}{k^0} \left( e^{-iky} - \theta (K_{max} - k^0) \right) \tilde{S}_{QCED}^{nls}$$
(7)

where  $K_{max}$  is a dummy parameter – here the same for QCD and QED.

#### Infrared Algebra(QCED):

$$\begin{split} x_{avg}(QED) &\cong \gamma(QED)/(1+\gamma(QED)) \\ x_{avg}(QCD) &\cong \gamma(QCD)/(1+\gamma(QCD)) \\ \gamma(A) &= \frac{2\alpha_A \mathcal{C}_A}{\pi} (L_s-1), A = QED, QCD \\ \mathcal{C}_A &= Q_f^2, C_F, \text{ respectively, for } A = QED, QCD \end{split}$$

 $\Rightarrow$  QCD dominant corrections happen an order of magnitude earlier than those for QED.  $_{\simeq(0,0)}$ 

 $\Rightarrow$  Leading  $\bar{\beta}_{0,0}^{(0,0)}$ -level gives a good estimate of the size of the effects we study.

**RESUMMED QUANTUM GRAVITY** 

APPLY (6) TO QUANTUM GENERAL RELATIVITY:

$$i\Delta_F'(k)|_{\text{resummed}} = \frac{ie^{B_g''(k)}}{(k^2 - m^2 - \Sigma_s' + i\epsilon)}$$
(8)

FOR

 $\Rightarrow$ 

$$B_g''(k) = -2i\kappa^2 k^4 \frac{\int d^4\ell}{16\pi^4} \frac{1}{\ell^2 - \lambda^2 + i\epsilon} \frac{1}{(\ell^2 + 2\ell k + \Delta + i\epsilon)^2}$$
(9)

THIS IS THE BASIC RESULT.

NOTE THE FOLLOWING:

•  $\Sigma_s'$  starts in  $\mathcal{O}(\kappa^2)$  , so we may drop it in calculating one-loop effects.

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• EXPLICIT EVALUATION GIVES, FOR THE DEEP UV REGIME,

$$B_g''(k) = \frac{\kappa^2 |k^2|}{8\pi^2} \ln\left(\frac{m^2}{m^2 + |k^2|}\right),$$
 (10)

 $\Rightarrow$  THE RESUMMED PROPAGATOR FALLS FASTER THAN ANY POWER OF  $|k^2|!$ 

• IF *m* VANISHES, USING THE USUAL  $-\mu^2$  NORMALIZATION POINT WE GET  $B''_g(k) = \frac{\kappa^2 |k^2|}{8\pi^2} \ln\left(\frac{\mu^2}{|k^2|}\right)$  WHICH AGAIN VANISHES FASTER THAN ANY POWER OF  $|k^2|$ !

THIS MEANS THAT ONE-LOOP CORRECTIONS ARE FINITE! INDEED, ALL QUANTUM GRAVITY LOOPS ARE UV FINITE(MPLA17(2002)2371)!



#### **QED QCD** Threshold Corrections, Shower/ME Matching

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and IR-Improved DGLAP Theory at LHC

We shall apply the new simultaneous QED $\otimes$ QCD exponentiation calculus to the sinlge Z production with leptonic decay at the LHC (and at FNAL) to focus on the ISR alone, for definiteness. See also the work of Baur *et al.*, Dittmaier and Kramer, Zykunov for exact  $\mathcal{O}(\alpha)$  results and Hamberg *et al.*, van Neerven and Matsuura and Anastasiou *et al.* for exact  $\mathcal{O}(\alpha_s^2)$  results.

For the basic formula

$$d\sigma_{exp}(pp \to V + X \to \bar{\ell}\ell' + X') = \sum_{i,j} \int dx_i dx_j F_i(x_i) F_j(x_j) d\hat{\sigma}_{exp}(x_i x_j s),$$
(11)

we use the result in (6) here with semi-analytical methods and structure functions from Martin *et al.*.

A MC realization will appear elsewhere.

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#### **SHOWER/ME MATCHING**

- Note the following: In (11) WE DO NOT ATTEMPT AT THIS TIME TO REPLACE HERWIG and/or PYTHIA – WE INTEND TO COMBINE OUR EXACT YFS CALCULUS,  $d\hat{\sigma}_{exp}(x_ix_js)$ , WITH HERWIG and/or PYTHIA BY USING THEM/IT "IN LIEU" OF  $\{F_i\}$ . A. USE HERWIG/PYTHIA SHOWER FOR  $p_T \leq \mu$ , YFS nG for  $p_T > \mu$ . B. EXPAND HERWIG/PYTHIA SHOWER FOR  $p_T \leq \mu$ , YFS nG for  $p_T > \mu$ . B. EXPAND HERWIG/PYTHIA SHOWER FORMULA $\otimes d\sigma_{exp}$  AND ADJUST  $\tilde{\beta}_{n,m}$  TO EXACTNESS FOR DESIRED ORDER WITH NEW  $\tilde{\beta}'_{n,m}$ FIRST USE  $\{F_i\}$  TO PICK  $(x_1, x_2)$ ; MAKE EVT WITH  $d\sigma_{exp}$ ; THEN SHOWER EVT USING HERWIG/PYTHIA VIA LES HOUCHES RECIPE.
- THIS COMBINATION OF THEORETICAL CONSTRUCTS CAN BE SYSTEMATICALLY IMPROVED WITH EXACT RESULTS ORDER-BY-ORDER IN  $\alpha_s, \alpha$ , WITH EXACT PHASE SPACE.
- THE RECENT ALTERNATIVE PARTON SHOWER ALGORITHM BY JADACH and SKRZYPEK, Acta. Phys. Pol.B35, 745 (2004), CAN ALSO BE USED.

• LACK OF COLOR COHERENCE  $\Rightarrow$  ISAJET NOT CONSIDERED HERE.

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With this said, we compute , with and without QED, the ratio

 $r_{exp} = \sigma_{exp} / \sigma_{Born}$ 

to get the results (We stress that we *do not* use the narrow resonance approximation here.)

 $r_{exp} = \begin{cases} 1.1901 &, \text{QCED} \equiv \text{QCD+QED}, \text{LHC} \\ 1.1872 &, \text{QCD}, \text{LHC} \\ 1.1911 &, \text{QCED} \equiv \text{QCD+QED}, \text{Tevatron} \\ 1.1879 &, \text{QCD}, \text{Tevatron} \end{cases}$ (12)

 $\Rightarrow$ 

**\*QED IS AT .3% AT BOTH LHC and FNAL.** 

**\*THIS IS STABLE UNDER SCALE VARIATIONS.** 

\*WE AGREE WITH BAUR ET AL., HAMBERG ET AL., van NEERVEN and

ZIJLSTRA.

\*QED EFFECT SIMILAR IN SIZE TO STR. FN. RESULTS.

**\* DGLAP SYNTHESIZATION HAS NOT COMPROMISED THE NORMALIZATION.** 

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IR-Improved DGLAP Theory

**APPLY QCD EXPN THEORY TO DGLAP KERNELS:** 

$$P_{qq}(z) = C_F F_{YFS}(\gamma_q) e^{\frac{1}{2}\delta_q} \left[ \frac{1+z^2}{1-z} (1-z)^{\gamma_q} - f_q(\gamma_q) \delta(1-z) \right]$$
(13)

where

$$f_q(\gamma_q) = \frac{2}{\gamma_q} - \frac{2}{\gamma_q + 1} + \frac{1}{\gamma_q + 2}$$
(14)

and

$$\gamma_q = C_F \frac{\alpha_s}{\pi} t = \frac{4C_F}{\beta_0} \tag{15}$$

$$\delta_q = \frac{\gamma_q}{2} + \frac{\alpha_s C_F}{\pi} (\frac{\pi^2}{3} - \frac{1}{2})$$
(16)

and

$$F_{YFS}(\gamma_q) = \frac{e^{-C_E \gamma_q}}{\Gamma(1 + \gamma_q)}.$$
(17)

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SIMILAR RESULTS HOLD FOR  $P_{Gq}, P_{GG}, P_{qG}$ , GIVING:

$$P_{qq}(z) = C_F F_{YFS}(\gamma_q) e^{\frac{1}{2}\delta_q} \left[ \frac{1+z^2}{1-z} (1-z)^{\gamma_q} - f_q(\gamma_q) \delta(1-z) \right],$$
(18)

$$P_{Gq}(z) = C_F F_{YFS}(\gamma_q) e^{\frac{1}{2}\delta_q} \frac{1 + (1 - z)^2}{z} z^{\gamma_q},$$

$$P_{GG}(z) = 2C_G F_{YFS}(\gamma_G) e^{\frac{1}{2}\delta_G} \{ \frac{1 - z}{z} z^{\gamma_G} + \frac{z}{1 - z} (1 - z)^{\gamma_G} + \frac{1}{2} (z^{1 + \gamma_G} (1 - z) + z(1 - z)^{1 + \gamma_G}) - f_G(\gamma_G) \delta(1 - z) \},$$
(20)

$$P_{qG}(z) = F_{YFS}(\gamma_G) e^{\frac{1}{2}\delta_G} \frac{1}{2} \{ z^2 (1-z)^{\gamma_G} + (1-z)^2 z^{\gamma_G} \},$$
(21)

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where

$$\gamma_G = C_G \frac{\alpha_s}{\pi} t = \frac{4C_G}{\beta_0} \tag{22}$$

$$\delta_G = \frac{\gamma_G}{2} + \frac{\alpha_s C_G}{\pi} (\frac{\pi^2}{3} - \frac{1}{2}), \tag{23}$$

$$f_G(\gamma_G) = \frac{n_f}{C_G} \frac{1}{(1+\gamma_G)(2+\gamma_G)(3+\gamma_G)} + \frac{2}{\gamma_G(1+\gamma_G)(2+\gamma_G)}$$
(24)  
+  $\frac{1}{(1+\gamma_G)(2+\gamma_G)} + \frac{1}{2(3+\gamma_G)(4+\gamma_G)}$ (25)  
+  $\frac{1}{(2+\gamma_G)(3+\gamma_G)(4+\gamma_G)}.$ (26)

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Parton Distributions

Moments of kernels  $\Leftrightarrow$  Logarithmic exponents for evolution

$$\frac{dM_n^{NS}(t)}{dt} = \frac{\alpha_s(t)}{2\pi} A_n^{NS} M_n^{NS}(t)$$
(27)

where

$$M_n^{NS}(t) = \int_0^1 dz z^{n-1} q^{NS}(z,t)$$
(28)

and the quantity  ${\cal A}_n^{NS}$  is given by

$$A_n^{NS} = \int_0^1 dz z^{n-1} P_{qq}(z),$$
  
=  $C_F F_{YFS}(\gamma_q) e^{\frac{1}{2}\delta_q} [B(n,\gamma_q) + B(n+2,\gamma_q) - f_q(\gamma_q)]$  (29)

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where B(x,y) is the beta function given by

$$B(x,y) = \Gamma(x)\Gamma(y)/\Gamma(x+y)$$

**Compare the usual result** 

$$A_n^{NS^o} \equiv C_F \left[ -\frac{1}{2} + \frac{1}{n(n+1)} - 2\sum_{j=2}^n \frac{1}{j} \right].$$
 (30)

- ASYMPTOTIC BEHAVIOR: IR-improved goes to a multiple of  $-f_q$ , consistent with  $\lim_{n\to\infty} z^{n-1} = 0$  for  $0 \le z < 1$ ; usual result diverges as  $-2C_F \ln n$ .
- Different for finite n as well: for n=2 we get, for example, for  $\alpha_s\cong .118$ ,

$$A_2^{NS} = \begin{cases} C_F(-1.33) &, \text{ un-IR-improved} \\ C_F(-0.966) &, \text{ IR-improved} \end{cases}$$
(31)

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#### • For completeness we note

$$M_{n}^{NS}(t) = M_{n}^{NS}(t_{0})e^{\int_{t_{0}}^{t} dt' \frac{\alpha_{s}(t')}{2\pi} A_{n}^{NS}(t')}$$
  
=  $M_{n}^{NS}(t_{0})e^{\bar{a}_{n}[Ei(\frac{1}{2}\delta_{1}\alpha_{s}(t_{0})) - Ei(\frac{1}{2}\delta_{1}\alpha_{s}(t))]}$  (32)

$$\implies_{t,t_0 \text{ large with } t >> t_0} M_n^{NS}(t_0) \left(\frac{\alpha_s(t_0)}{\alpha_s(t)}\right)^{a'n}$$

where  $Ei(x) = \int_{-\infty}^{x} dr e^r / r$  is the exponential integral function,

$$\bar{a}_{n} = \frac{2C_{F}}{\beta_{0}} F_{YFS}(\gamma_{q}) e^{\frac{\gamma_{q}}{4}} [B(n,\gamma_{q}) + B(n+2,\gamma_{q}) - f_{q}(\gamma_{q})]$$

$$\bar{a'}_{n} = \bar{a}_{n} \left( 1 + \frac{\delta_{1}}{2} \frac{(\alpha_{s}(t_{0}) - \alpha_{s}(t))}{\ln(\alpha_{s}(t_{0})/\alpha_{s}(t))} \right)$$
(33)

with

$$\begin{split} \delta_1 &= \frac{C_F}{\pi} \left( \frac{\pi^2}{3} - \frac{1}{2} \right) \\ \text{Compare with un-IR-improved result where last line in eq.(32) holds exactly with} \\ \bar{a'}_n &= 2A_n^{NS^o} / \beta_0. \end{split}$$

• Comparison with Moch et al., Vogt et al., etc., in progress.

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FINAL STATE OF HAWKING RADIATION

CONSIDER THE GRAVITON PROPAGATOR IN THE THEORY OF GRAVITY COUPLED TO A MASSIVE SCALAR(HIGGS) FIELD(Feynman). WE HAVE THE GRAPHS



Figure 1: The graviton((a),(b)) and its ghost((c)) one-loop contributions to the

graviton propagator. q is the 4-momentum of the graviton.

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Figure 2: The scalar one-loop contribution to the graviton propagator.  $\boldsymbol{q}$  is the

4-momentum of the graviton.



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USING THE RESUMMED THEORY, WE GET THAT THE NEWTON POTENTIAL BECOMES

$$\Phi_N(r) = -\frac{G_N M}{r} (1 - e^{-ar}), \tag{34}$$

FOR

$$a \cong 0.210 M_{Pl}.\tag{35}$$

CONTACT WITH AYMPTOTIC SAFETY APPROACH

• OUR RESULTS IMPLY

$$G(k) = G_N / (1 + \frac{k^2}{a^2})$$

 $\Rightarrow$  FIXED POINT BEHAVIOR FOR

 $k^2 \rightarrow \infty$ ,

IN AGREEMENT WITH THE PHENOMENOLOGICAL ASYMPTOTIC SAFETY APPROACH OF BONNANNO & REUTER IN PRD62(2000) 043008.





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A FURTHER "AGREEMENT": FINAL STATE OF HAWKING RADIATION OF AN **ORIGINALLY VERY MASSIVE BLACKHOLE BECAUSE OUR VALUE OF THE COEFFICIENT,**  $\frac{1}{a^2}$ , OF  $k^2$  IN THE DENOMINATOR OF G(k)AGREES WITH THAT FOUND BY BONNANNO & REUTER(B-R), IF WE USE THEIR PRESCRIPTION FOR THE RELATIONSHIP BETWEEN k and rIN THE REGIME WHERE THE LAPSE FUNCTION VANISHES, WE GET THE SAME HAWKING RADIATION PHENOMEMNOLOGY AS THEY DO: THE BLACK HOLE EVAPORATES IN THE B-R ANALYSIS UNTIL IT REACHES A MASS

 $M_{cr} \sim M_{Pl}$ 

AT WHICH THE BEKENSTEIN-HAWKING TEMPERATURE VANISHES, LEAVING A PLANCK SCALE REMNANT.

• FATE OF REMNANT? IN hep-ph/0503189  $\Rightarrow$  OUR QUANTUM LOOP EFFECTS COMBINED WITH THE G(r) OF B-R IMPLY HORIZON OF THE PLANCK SCALE REMNANT IS OBVIATED – CONSISTENT WITH RECENT RESULTS OF HAWKING.

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#### TO WIT, IN THE METRIC CLASS

$$ds^{2} = f(r)dt^{2} - f(r)^{-1}dr^{2} - r^{2}d\Omega^{2}$$
(36)

THE LAPSE FUNCTION IS, FROM B-R,

$$f(r) = 1 - \frac{2G(r)M}{r} = \frac{B(x)}{B(x) + 2x^2} |_{x = \frac{r}{G_N M}},$$
(37)

**WHERE** 

$$B(x) = x^3 - 2x^2 + \Omega x + \gamma \Omega \tag{38}$$

FOR

$$\Omega = \frac{\tilde{\omega}}{G_N M^2} = \frac{\tilde{\omega} M_{Pl}^2}{M^2}.$$
(39)



AFTER H-RADIATING TO REGIME NEAR  $M_{cr} \sim M_{Pl}$ , quantum loops allow US to replace G(r) with  $G_N(1 - e^{-ar})$  in the lapse function for  $r < r_>$ , the outermost solution of

$$G(r) = G_N(1 - e^{-ar}).$$
 (40)

IN THIS WAY, WE SEE THAT THE INNER HORIZON MOVES TO NEGATIVE r and the outer horizon moves to r=0 at the new critical mass  $\sim 2.38 M_{Pl}.$ 

NOTE: M. BOJOWALD *et al.*, gr-qc/0503041, – LOOP QG CONCURS WITH GENERAL CONCLUSION.

PREDICTION: THERE SHOULD ENERGETIC COSMIC RAYS AT  $E\sim M_{Pl}$  due the decay of such a remnant.

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YFS THEORY ( EEX AND CEEX) EXTENDS TO NON-ABELIAN GAUGE THEORY AND ALLOWS SIMULTANEOUS EXPN OF QED AND QCD WITH PROPER SHOWER/ME MATCHING BUILT-IN. FOR QED © QCD

- FULL MC EVENT GENERATOR REALIZATION IS POSSIBLE.
- SEMI-ANALYTICAL RESULTS FOR QED (AND QCD) THRESHOLD EFFECTS AGREE WITH LITERATURE ON Z PRODUCTION
- AS QED IS AT THE .3% LEVEL, IT IS NEEDED FOR 1% LHC THEORY PREDICTIONS.
- A FIRM BASIS FOR THE COMPLETE  $\mathcal{O}(\alpha_s^2, \alpha \alpha_s, \alpha^2)$  MC RESULTS NEEDED FOR THE FNAL/LHC/RHIC/TESLA/LC PHYSICS HAS BEEN DEMONSTRATED AND ALL THE LATTER IS IN PROGRESS.

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THE THEORY ALLOWS A NEW APPROACH TO QUANTUM GENERAL RELATIVITY: • RESUMMED QG UV FINITE • MANY CONSEQUENCES: BLACK HOLES EVAPORATE TO FINAL MASS  $\sim M_{Pl}$ WITH NO HORIZON

 $\Rightarrow E \sim M_{Pl}$  COSMIC RAYS, · · · .

