

Magnetic moment of the baryon at finite density and temperature

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- Introduction
- Model
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- Conclusion

Hyp-X 2009, Tokai
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INTRODUCTION

- High density and/or temperature
 - Nuclear interactions
 - Hadron properties
 - Symmetry restoration
 - Phase transition
- Probing tool
 - Neutron star observation
 - Relativistic heavy ion collisions

- Magnetic form factor in nuclei : finite density
 - Recent experiment at JLab ~ 5% change in nuclei

S. Strauch, arXiv:0709.4034v1 [nucl-ex]

- Yet, old data are controversial

K.S.Kim, M.K. Cheoun, PRC67, 034603 (2003)

- Heavy ion collision : finite temperature
 - Strong magnetic field in peripheral collision $\sim 10^{18}$ G

D.E. Kharzeev, L.D. McLerran, H.J. Warringer, arXiv:0711.0950 [hep-ph]

- Strong magnetic field can change the properties of the baryon in free space

B.C. Tiburzi, NPA814, 74 (2009)

- Strong magnetic field can change the critical density and temperature for the phase transition

N.O. Agasian, S.M. Fedorov, PLB663, 445 (2008)

- Magnetic moment of the baryon at high density and temperature
 - Probe the change of hadron properties

MODEL

- Quark-meson coupling model
 - Baryons : MIT bags
 - Interactions : Mean fields of σ and ω mesons

$$[i\gamma \cdot \partial - (m_q - g_\sigma^q \sigma) - g_\omega^q \gamma^0 \omega_0] \psi_q = 0$$

$$m_B^* = \sqrt{E_B^2 - \sum_q \left(\frac{x_q}{R_B}\right)^2}$$

$$E_B = \sum_q \frac{\Omega_q}{R_B} - \frac{Z_b}{R_B} + \frac{4\pi R_B^3}{3} B_B$$

- Parameters

- Mass in free space : Bag constants (B_B),

Phenomenological constants (Z_b)

- Nuclear saturation properties : quark-meson coupling constants (g_σ^q, g_ω^q)

- QMC $B_B = B_{B0}$

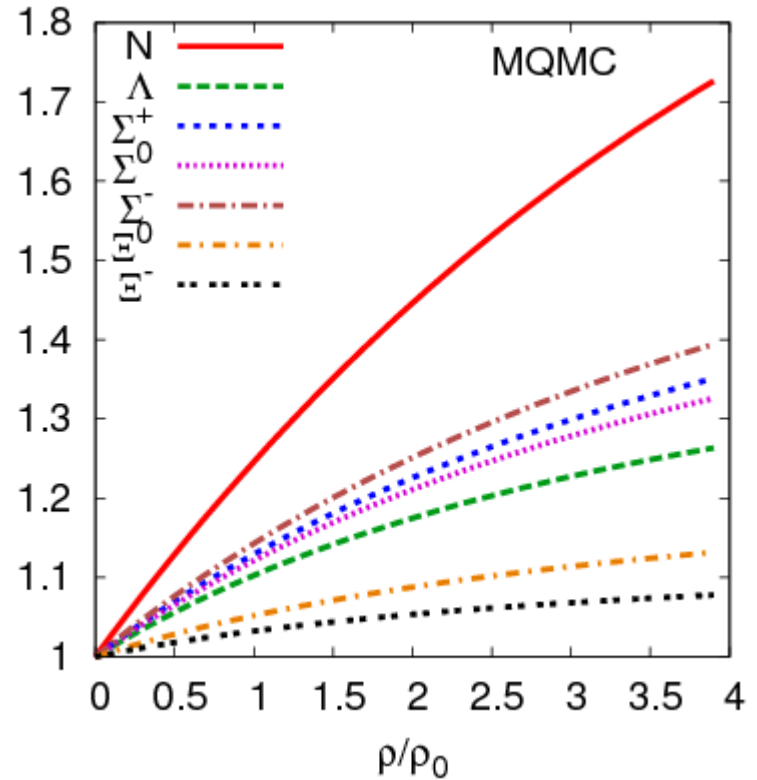
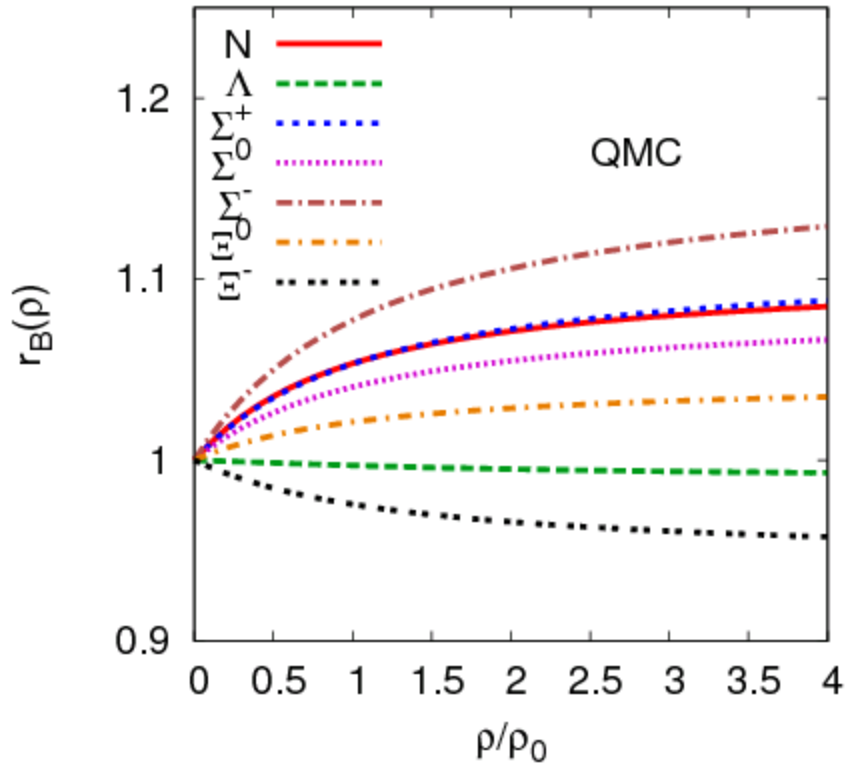
- MQMC $B_B = B_B(\sigma)$

$$B_B(\sigma) = B_{B0} \exp\left(-\frac{4g_\sigma^B \sigma}{M_N}\right)$$

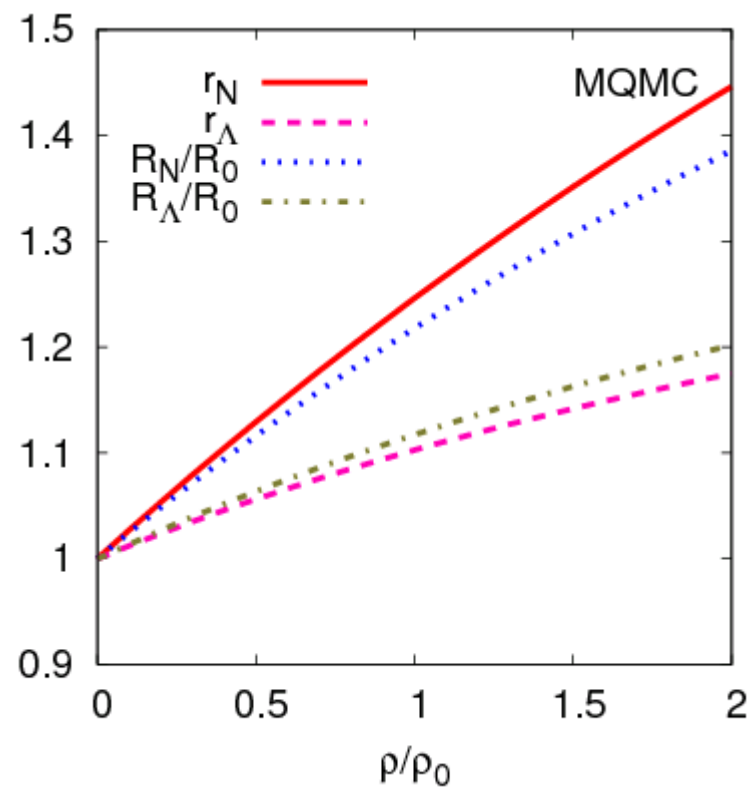
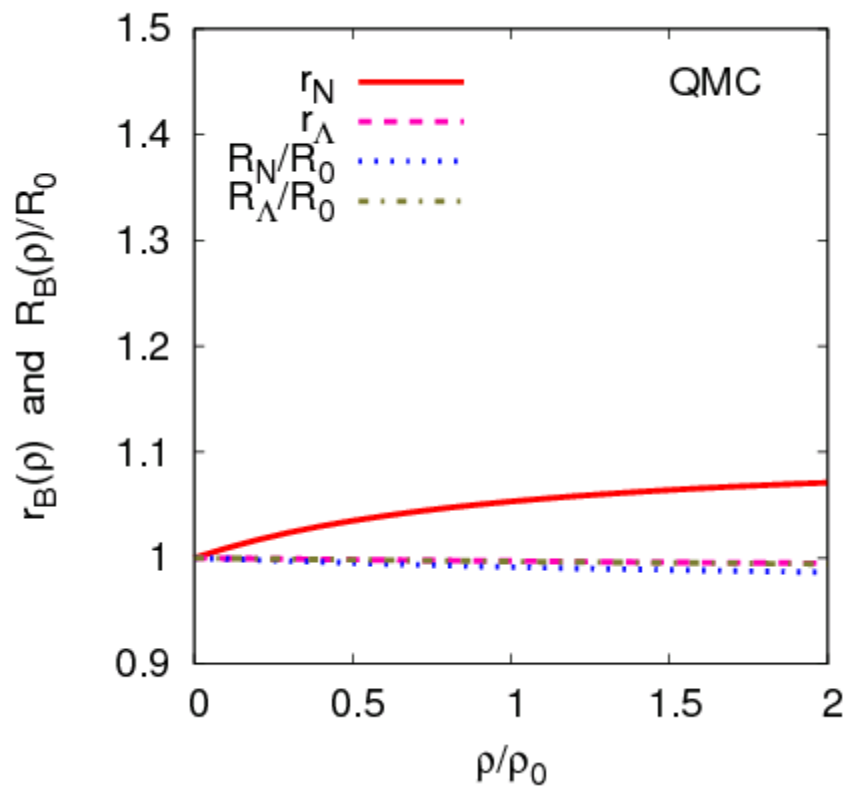
RESULTS

Magnetic moment ratio to free space value : $r_B = \mu_B^*/\mu_B$

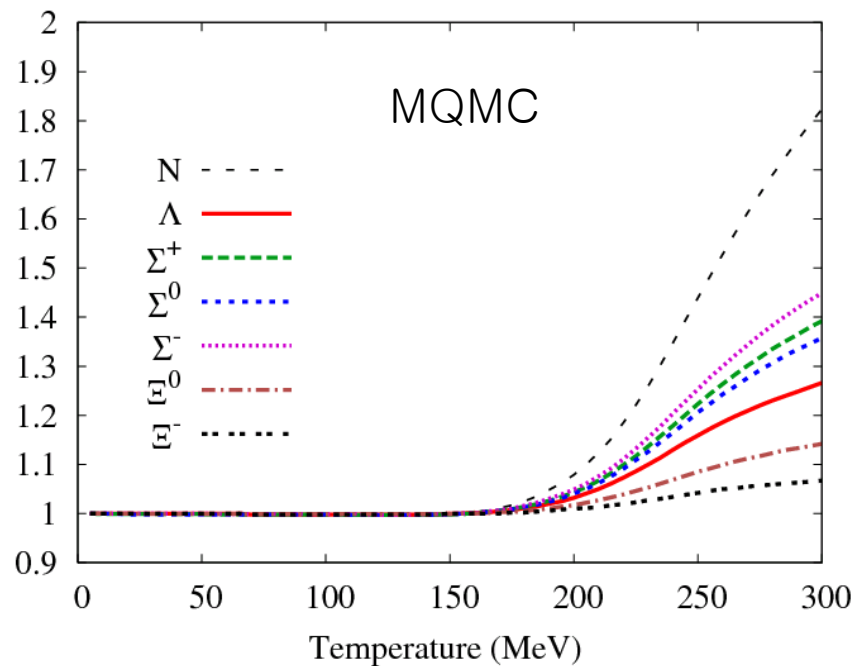
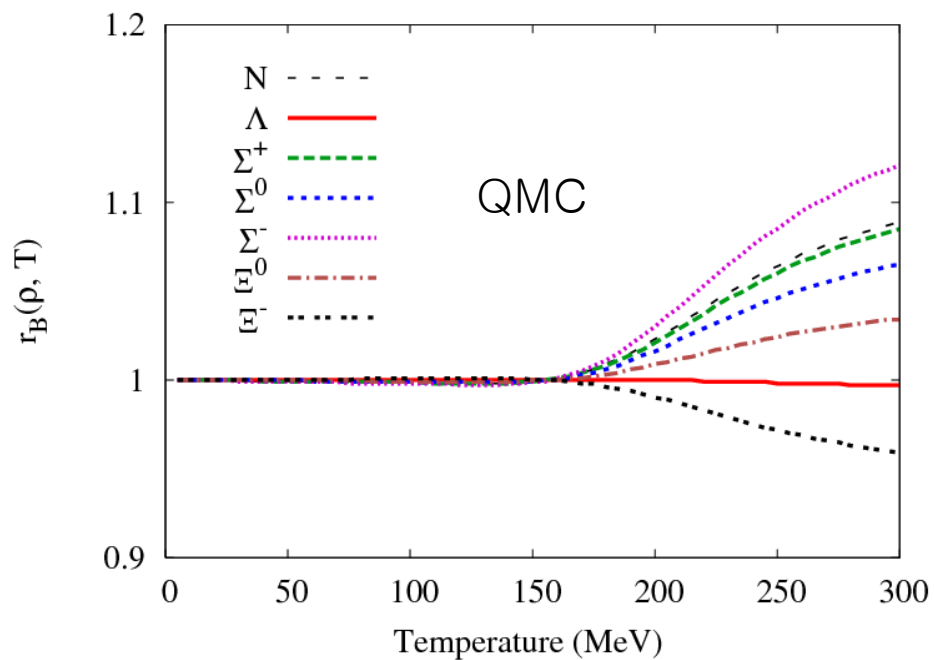
- Finite density, $T=0$



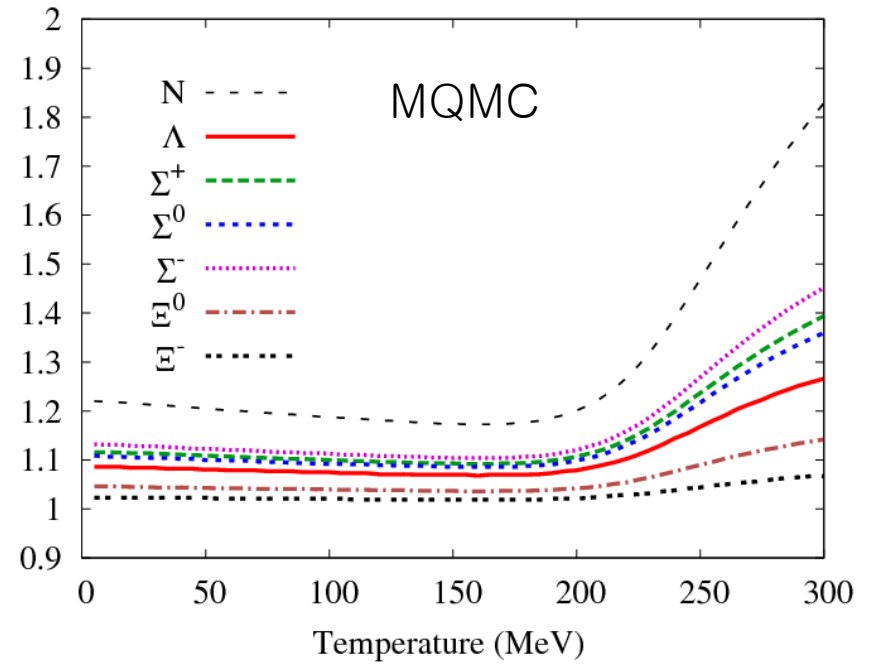
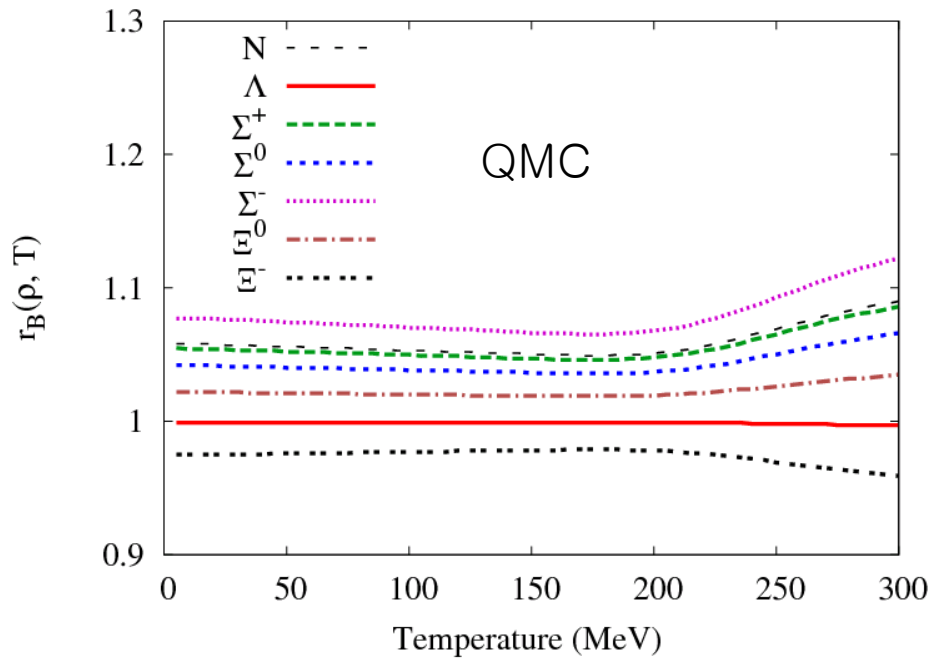
Explanation for the model dependence : $\mu_B \sim R_B$



- Finite temperature, $\rho=0.1\rho_0$



- Finite temperature, $\rho=1.0\rho_0$



- The effect observable?

$$E_N = m_N - \mu_N eB/2m_N$$

$(eB)^{1/2} \sim 10^{18}$ G in peripheral heavy ion collision

- In free space

$$\mu_p eB/2m_N \sim \mathbf{3\% \text{ of } m_N}$$

- For $T=250$ MeV and $\rho=0.1\rho_0$ in the MQMC model

$$m_N^* \sim 0.7 m_N, \mu_N^* \sim 1.5 \mu_N$$

$$\mu_p^* eB/2m_N^* \sim \mathbf{10\% \text{ of } m_N^*}$$

The effect is obvious

CONCLUSION

- Obvious change at finite density and temperature
- Density effect
 - Neutron star phenomenology in combination with its strong magnetic field
- Temperature effect
 - Baryon mass shift in relativistic heavy ion collisions
- Further investigations
 - Relation to symmetry restoration
 - Critical density and temperature for phase transition
 - Magnetic moment at high density in the strong magnetic field in progress