Vector meson mass modification and its effect to dense matter in RMF model

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Introduction

Interest: dense matter EOS including hyperon



Flavor symmetry, Hypernuclei

Partitional restoration at high $\rho_{\rm B}$

[1] Lattimer and Swesty, NPA 535 (1991) 331.
[2] Shen et al., PTP 100 1013.
[3] Yamamoto, Nishizaki and Takatsuka, NPA 691 (2001) 432.
[4] Ogawa et al., PTP 111 (2004) 75.

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Chiral symmetric RMF

Strong point of RMF

- Good description of Nuclei and Nuclear Matter
- Symmetries are naturally included
- Problems in chiral RMF [5]
 - 1. Sudden chiral restoration below ρ_0 [6]
 - 2. Too stiff EOS [4, 7]
 - 3. Instability at large chiral condensate

[5] KT and Ohnishi, Prog. Theor. Phys. 117 (2007) 903. [6] Matsui and Serot, AP 144 (1982) 107.[4] Ogawa et al, PTP 111 (2004) 75. [7] Boguta, PLB 120 (1983) 34.

Energy Density at $\rho_{\rm B}$ =(0-5) $\rho_{\rm 0}$

Effective Potential in SCL-LQCD (Zero T)

QCD Lattice Action (Zero T treatment)

Kawamoto, Smit, 1981 $S = S_G + S_F + m_0 \overline{X} X$ Strong Coupling Limit $\rightarrow -\frac{1}{2}(\overline{X}X)V_M(\overline{X}X) + m_0\overline{X}X$ **One-link integral** (1/d expansion*) $\rightarrow \frac{1}{2} \sigma V_M^{-1} \sigma + \overline{\chi} (\sigma + m_0) \chi$ **Bosonization** $\rightarrow \frac{1}{2} \sigma V_M^{-1} \sigma - N_c \sum \log(\sigma(x) + m_0)$ Fermion Integral $= L^{d} N_{\tau} \left| \frac{N_{c}}{d+1} \bar{\sigma}^{2} - N_{c} \log(\bar{\sigma} + m_{0}) \right|$

* d = Spatial dim.

Effective Potential

Fermion Matrix = Just a number \rightarrow Simple Logarithmic Effective Potential for σ

$$V_{\sigma} = \frac{1}{2} a_{\sigma} \sigma^2 - b_{\sigma} \log \sigma$$

Ohnishi, YITP Colloq., 2008/05/28

Chiral potential from SCL

 Chiral potential: derived from Strong Coupling Limit of lattice QCD [8,9]

$$U_{\sigma} = -a \log(\det M_{SU(2)}M^{\dagger}_{SU(2)}) + btr(M_{SU(2)}M^{\dagger}_{SU(2)}) + c_{\sigma}\sigma$$

$$= -2a\left\{\log\left(1 + \frac{\sigma}{f_{\pi}}\right) + \frac{\sigma}{f_{\pi}} - \frac{1}{2}\left(\frac{\sigma}{f_{\pi}}\right)^{2}\right\} + \frac{1}{2}m_{\sigma}^{2}\sigma^{2} + \frac{1}{2}m_{\pi}^{2}\pi^{2}$$

[8] Kawamoto and Smit, NPB 192 (1981) 100. [9] KT and Ohnishi, PTP 117 (2007) 903. [10] Sugahara and Toki, NPA 579 (1994) 557.

Λ and **Σ** hypernuclear results

Results of NS matter (1)

- Extra degree of freedom: softening Sym. Matter EOS.
- Softer than TM model [8] or SU(2) version especially in high density.
- E/B in RMF is almost linear at high densities cf. ρ_B² in FP

Results of NS matter (2)

 Max Mass of NS in chiral RMF =1.3M with parameter sets determined from hypernuclear data 2.5 Masses of hadronic stars \rightarrow Underestimate 2 $M_{max}/M_{sun}=2.23$ $M_{max}/M_{sun}=1.65$ existing NS mass ! M/M_{sun} 1.5 $M_{max}/M_{sun}=1.34$ SCL3 1 SCL3/ TM1 0.5 0

We need for extra repulsion at high $\rho_{\rm B}$ as pointed by Takatsuka

0.5

0

1.5

1

 $\rho_B~(\text{fm}^{-3})$

2

Mass modification

- Sym. nuclear matter EOS: good agreement with FP EOS around ρ₀
 - This EOS is too soft at high $\rho_{\rm B}$ to support NS mass.
 - Origin of repulsion: partitional σω coupling introduced to reproduce ω mass reduction[11]?

$$\frac{1}{2}m_{\omega}^{2}(\sigma)\omega^{2} = \frac{1}{2}\left[(m_{\omega}^{0})^{2} + g_{\sigma\omega}(f_{\pi} - \sigma)^{2}\right]\omega^{2}$$

[11] Naruki et al., PRL 96 (2006) 092301.

Effect of mass modification to dense matter in RMF model

- Suggested mass reduction of ω: 9%
- Sym. nuclear matter EOS becomes softer after fitting BE of nuclei !
- Why ? Smaller ω mass at ρ_0 \rightarrow smaller $g_{\omega N}$ \rightarrow softer EOS at high ρ_{R}
- Do we miss something ?

BE with modified meson mass

- With medium modification of vector meson mass, we cannot describe binding energies of light nuclei (<Ca).
- Similar results are obtained by Furnstahl et al. and Heide et al. with glueball field representing scale anomaly.
 Similar results are obtained by Furnstahl et al. Lighter side 9 8 7 6

Summary

- We have invesitgated the properties of normal nuclei and hypernulei in chiral SU(3) RMF.
 - Energy density at $\rho_B = 0 \rightarrow SCL-LQCD$
 - Vector meson mass: Constant / Modified at finite ρ_B
- Properties of sym. nuclear matter, normal nuclei, Λ and Σ hypernuclear state are well reproduced with constant ω mass, but max. NS mass is calculated to be too small (1.3M_o).
- σω coupling: medium mod. of vector meson mass may not lead to the universal extra repulsion.

Thank you for attention!! and KT is so sorry not to be here!!

A short-cut from QCD to Nuclei

Ohnishi, YITP Collog., 2008/05/28