Prospect of neutrino interaction and nuclear effect studies

Yoshinari Hayato (Kamioka Obs., ICRR, U-Tokyo)

Introduction



Neutrino-nucleon interactions from 100 MeV ~ a few GeV

Charged current quasi-elastic scattering (CCQE)

 $v + n \rightarrow l + p$

Neutral current elastic scattering

 $v + N \rightarrow v + N$

Single meson productions

 $v + N \rightarrow I + N' + \pi (\eta, K)$

Single photon productions (radiative decay of resonance)

 $\nu + N \rightarrow I + N' + \gamma$

Deep (/ shallow) inelastic scattering $v + N \rightarrow I + N' + m\pi(\eta, K)$





Neutrino-nucleus interactions from 100 MeV ~ a few GeV

Most of the neutrino interaction targets are in nucleus. Different from the 'free' nucleon. V

Target nucleon is in the nuclear medium (matter) Impulse approximation and Fermi-gas model works fairly well in this energy region.

Interacted nucleon or produced mesons interact with the other nucleons in nucleus, so-called final state interactions.

However, this simple models (treatments) are not precise enough for the recent experiments.





Charged current quasi-elastic scattering $v + n \rightarrow l + p$

Current (experimental) situation

- Low energy experiments (K2K-SciFi, MiniBooNE, E_v < a few GeV) 1) small q² or very forward region seems to have larger suppression than naive model. 2) interaction probability (cross-section) seems to be larger than expected.
- Medium energy experiment (MINERvA, $< E_v > ~ 3.5$ GeV), discrepancies seem to be smaller.
- High energy experiment (NOMAD, $< E_v > ~ 24 \text{ GeV}$)

Consistent with the simple Fermi-Gas model...



Charged current quasi-elastic scattering v + n → t + p

Several possible improvements have been proposed.

• Random phase approximation (RPA) correction

the presence of strongly interacting nucleons in nucleus may change the he weak interaction strength

Multi-nucleon effects

 Interactions with strongly
 coupled nucleons.

Confirmed w/ electron scat. exp.





Charged current quasi-elastic scattering v + n →I⁺ + p ~ Fit existing data sets with new models ~

a) Relativistic Fermi Gas model with RPA correction (Nieves et al.)
+ Multi-Nucleon interactions (Nieves et al.)
b) Spectral function model (A. Ankowski, O. Benhar et al.)

+ Multi-Nucleon interactions (Nieves et al.)





* Additional normalization factors are included for MiniBooNE data

Selection of the model based on the fit is difficult at this moment. Also, some tensions between data sets were observed.

For the oscillation analyses, we performed the fit with ND280 data to determine/ constraint the neutrino interaction parameters and neutrino flux simultaneously.

3 sub-samples from neutrino running

- Events with μ^- and no pion (CC0 π) CCQE enriched
- Events with μ^- and 1 π^{\pm} (CC1 π) CC 1 π enriched
- Other events with μ- (CC Other)
 CCDIS enriched





For the oscillation analyses, we performed the fit with ND280 data to determine/ constraint the neutrino interaction parameters and neutrino flux simultaneously.

2 sub-samples of $\overline{\nu}$ in anti-neutrino running period.

- Events with $\mu^{\text{+}}$ and only 1 track reconstructed CCQE enriched
- Events with μ^+ and there are more than 1 track reconstructed Inclusive (less CCQE)



Neutrino interaction and flux parameter fit using the ND280 data sets

Parameter <i>preliminary</i>	Prior	Fit result
CCQE M_A (GeV/c ²)	1.15 ± 0.07	1.13 ± 0.03
PF (Carbon) (MeV/c)	223.0 ± 12.3	222.7 ± 8.83
MEC fraction (%)	27.0 ± 29.0	103.1 ± 17.2

In the fit, not only these parameters but other parameters including flux were fit simultaneously.

Therefore,

these values can not be compared directly with the results from external data fits.

T2K ND280 data sets

seem to require MEC-like interactions. Need further dedicated studies!



Neutrino-nucleus interaction studies in T2K Cross-section measurements with the near detectors

• We have studied CCQE using off-axis detector. (arXiv: 1411.6264) Obtained (effective) $M_A = 1.26^{+0.21}_{-0.18}$ GeV/c² (absolute).

(without recent multi-nucleon interaction models.)

- Further detailed study of CCQE and Multi-nucleon interactions are required.
 - Now we are trying to use proton track information and compare with recent models with additional interactions.



Cross-section measurements with the near detectors

Data sets are compared with two model predictions:

Martini et al. RPA + Multi-nucleon interactions

Phys. Rev. C 80 065501 (2009)

---- Nieves et al. RPA + Multi-nucleon interactions

Phys. Rev. C 83 045501 (2011)



Cross-section measurements with the near detectors

We also studied CCQE using proton track with T2K-INGRID detector.



Single track (μ^- only) and Two track (μ^- + p) samples gave slightly different results.

Incompleteness of the model? ¹⁰⁻¹
Interesting to study the energy dependence using both on-axis and off-axis detector data sets.
Activities (~ energy deposit) around the vertex is

also expected to provide additional information.

(on-axis ~ higher energy region)



Cross-section measurements with the near detectors

On-Axis detector (INGRID) has scintillator module

INGRID

Proton Module

and Iron-scintillator composite modules.

Inclusive cross-section measurement

 $\frac{\sigma_{CC}^{Fe}}{\sigma_{CC}^{CH}} = 1.047 \pm 0.007 (stat.) + 0.035 (syst.)$ Phys. Rev. D 90, 052010 (2014)

Fe/C inclusive cross-section ~ 5% difference



Cross-section measurements with the near detectors Study of nuclear dependence ~ Use passive water layers in ND280

(One example of the on-going study ~ CC π^+ production)



We also working on the CCQE (-like) interactions.

Nuclear effects ~ pion interaction in nucleus

Pion interactions in nucleus is also important because these interactions affect determination of interaction mode.



Nuclear effects ~ pion interaction in nucleus

Recently DUET experiments measured

absorption + charge-exchange cross-sections.



Study to separate charge exchange from absorption is on-going.

1) **CCQE and Multi-nucleon CCQE-like interactions**

1. Measured interaction rates: MiniBooNE found larger interaction rate of CCQE(+like) interactions. Atmospheric sub-GeV v_{ρ} event rate seems to be consistent with $M_{A}=1.2 \text{GeV/c}^{2}$. (simple relativistic Fermi-Gas model) But there are uncertainties of v fluxes. 2. Suppressions in the small q^2 region / forward going leptons have been observed in various experiments: K2K, MiniBooNE, MINOS (preliminary), T2K.

Larger M_A (with simple Fermi-Gas model) seems to reproduce these to some extent.

1) **CCQE and Multi-nucleon CCQE-like interactions.**

Even though, low energy data (somewhat) agrees with expectations with simple Fermi-Gas + larger M_A, it is far from satisfactory.

Because possible contributions are neglected in the model:

RPA, multi-Nucleon interactions, Local density effects (local Fermi-Gas) or Spectral function (momentum-potential distributions) etc...

And assumed true E_v and reconstructed E_v relation are incorrect for Multi-nucleon interactions.

However, current data sets are not sufficient to justify or reject the models.

1) CCQE and Multi-nucleon CCQE-like interactions.

1) Further improvements of reconstruction tools and analyses (in T2K ND280) are essential.

Example)

Extensive use of the proton tracks and the energy deposit around the vertex

will provide useful information.

(Some indications in MINERvA results.)

 2) More neutrino and anti-neutrino data (of the T2K ND280) Still statistically limited to perform detailed analyses. Neutrino and anti-neutrino interactions have different dq²/dσ. This may provide additional information on nuclear effects and dependences.

1) CCQE and Multi-nucleon CCQE-like interactions.

With the current T2K ND280,

energy/momentum threshold (from track ID) or energy resolution (from vertex activity) of protons may not be sufficient to give definite answer. Also, we need to understand the nuclear dependence, i.e. difference between Carbon and Oxygen. (Current ND280 can provide some info. but may not be sufficient.) Therefore, we need additional information. Experiments like MINERvA and MicroBooNE will provide useful information. However, the current issues seem to have energy dependence. Also, Ar may be too large nucleus, unfortunately.

We need a sets of new detectors!

1) **CCQE and Multi-nucleon CCQE-like interactions.**

Several new detectors have been proposed. Among of them, a) Water + Scintillator 3D grid tracking detector (WAGASCI)

- b) Introducing water based scintillator cells in FGD
- c) Emulsion detector with water target
- d) High pressure gas TPC with Neon
 - seem to be quite interesting and promising.
- a) and b) are quite straight forward
 - and will improve understanding of C/O differences.
- c) and d) are really attractive
 - because they have low momentum threshold

with fine tracking capabilities.

23

d) may need (quite a lot) of R&D, though.

~ New detectors ~ WAGASCI(T59)



Emulsion detector(T60)





Neutrino-nucleus interaction studies to maximize neutrino oscillation sensitivity

(personal thoughts)

0.6

CCQE Cross-section (nucleon in Oxygen)

0.4

 $v_{\mu}/v_{e}\sigma$ ratio

2) v_{μ}/v_e ratio

Low energy σ (< 0.5 GeV) has nuclear model dependence. It may not have large impact but need to be studied.

3) Single π production cross-section and π momentum distribution $_{0}$

Gev Rein-Sehgal (FKR) model is known to have problem in reproducing electro-pion production data. Modified (improved) structure function is used in Neut but it is time to try more sophisticated models for comparison. (At this moment, there seems no significant difference, though.)

σ (10⁻³⁸cm²

0.5

0.2

 v_{e}

0.8

Neutrino-nucleus interaction studies to maximize neutrino oscillation sensitivity

(personal thoughts)

4) π interactions in/with Carbon and Oxygen

DUET experiment is expected to provide cross-sections of π + C scattering. However, p_{π} < 300 MeV/c.

We need to have higher P_{π} data for the improvements of

a) π secondary interactions (FSI) in nucleus

and b) π interactions in the detectors.

GEANT4 is used to simulate detector responses. Precisions of π interactions in the detector are important to study neutrino intearctions.



5) *nucleon interactions in/with Carbon and Oxygen* We need to find the way to improve nucleon interactions in the nucleus or in the detector.

6) Improvements of the neutrino interaction simulation programs

Recently, collaborations with theorists are really working well to improve our understanding of the models and the simulation software.

We have to continue and expand these activities not only for the neutrino oscillation studies but also for the better understanding of the neutrino-nucleus interactions.

Summary

Uncertainties of

- neutrino-nucleus interactions,
- hadron interactions in nucleus and
- hadron interactions in detectors are expected to be the largest sources of the systematic errors in the oscillation analyses.

 We have to improve analysis tools of ND280 for further detailed analyses of various interactions.
 Already running or starting experiments in the world are also expected to provide useful information.
 However, we also need `optimized' detectors with sufficient statistics

for better understandings of these interactions in the T2K energy region.

Summary

Recently, we are collaborating with theorists to understand the interactions and it is really helpful. We have to expand this kind of activities to develop more accurate and reliable neutrino-nucleus and detector simulation programs and to assign appropriate systematic uncertainties.

All the studies of these interactions will help in improving the atmospheric neutrino oscillation analyses and proton decay search sensitivities.

fin.

backups and memos

Study of neutrino interactions

Past experiments Bubble chamber ~ D₂ target (quasi-) free neutron data as reference

Recent experiments

MiniBooNE

Mineral Oil Cherenkov detector

 4π coverage

MINERnA

Fine grained tracking detector
Scintillator + other target material
Limited acceptance (< ~ 20 degree for μ etc.)
but sensitive to the heavy particles.

Charged current quasi-elastic scattering $v + n \rightarrow l + p$

Vector & Axial form factors

Vector part : Determined using electron scattering data

Axial part

: Determined using neutrino scattering data (dipole form, parameter is M_A)



The K2K experiment

Comparison with K2K 1kt detector and simulation data

Direction of μ w.r.t. beam events / 10 degrees 3500 Forward deficit - Data 3000 — NEUT-0000 — NEUT-1211 2500 2000 slightly larger *# of events* 1500 in backward 1000 500 0 50 150 100 θ_.. (degree) **NEUT-1211** $M_A = 1.1 \text{ GeV/c}^2$ for CCQE M_{Δ} =1.2 GeV/c² for 1 π

GRV94 with Bodek-Yang corr.

Coherent π by Marteau et.al.

Fully contained 1 ring μ -like events

NEUT-0000 $M_A=1.0 \text{ GeV/c}^2 \text{ for CCQE}$ $M_A=1.0 \text{ GeV/c}^2 \text{ for 1p}$ GRV94, Rein Sehgal coherent p Less forward going μ More backward going μ Agreements was better if we increased $M_{A.}$

Also, need corrections on
 DIS parton distribution function
 (Suggested by Bodek and Yang)
 and suppression of CC Coherent π
 production.

Important to measure large angle scattering!

Charged current Quasi-elastic scattering



Charged current Quasi-elastic scattering

MiniBooNE

Axial vector form factor parameter M_A



 $M_{\text{A}}\text{=}1.35\pm0.17~\text{GeV}$

 $F_{A}(Q^{2}) = \frac{g_{A}}{(1+Q^{2}/M_{A}^{2})^{2}}$

Need to be determined from the neutrino scattering data.

 ν_{μ}

n

- World avg. $M_{\text{A}}\text{=}1.02\pm0.17~\text{GeV}$
- **K2K SciFi** (¹⁶O, Q²>0.2) Phys. Rev. **D74**, 052002 (2006) M_A=1.20 ± 0.12 GeV
- K2K SciBar (¹²C, Q²>0.2) M_A =1.14 ± 0.11 GeV

 μ^{-}

р

 W^+

Charged current quasi-elastic scattering

NOMAD experiment

Carbon target, $\langle E_v \rangle \sim 24 \text{ GeV}$ $M_A(v) = \mathbf{1.06} \pm 0.02(stat.)$ $\pm 0.06 (syst.) \text{ GeV/c}^2$ $M_A(\bar{v}) = \mathbf{1.06} \pm 0.07(stat.)$ $\pm 0.10 (syst.) \text{ GeV/c}^2$



37

Consistent with Bubble chamber results.



Charged current quasi-elastic scattering $v + n \rightarrow l^{+} + p$ MINERvA Experiment $\langle E_v \rangle \sim 3.5$ GeV



Simple relativistic Fermi Gas model $M_A = 0.99 \text{ GeV/c}^2$

did not give the best fit result. Need some modification

